

# Plasma environment of Europa

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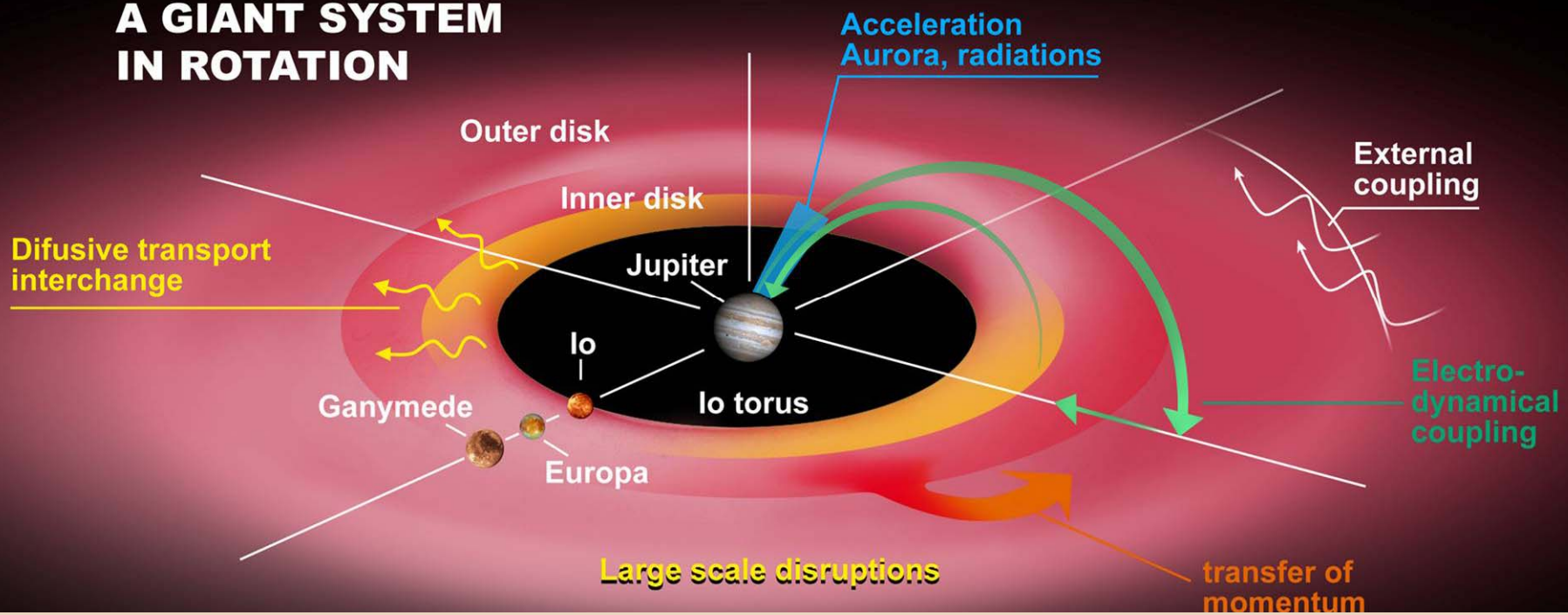
a review of plasma conditions and effects near Europa  
of interest to prime tasks of a landing mission

**Acknowledgements** to used publications:

K.Khurana et al, Russell et al, Frank et al., Paranicas et al., Kivelson et al.,  
Johnson et al., Krupp et al., etc



# A GIANT SYSTEM IN ROTATION



## Jupiter specifics

weak solar wind: ★ internal effects dominate

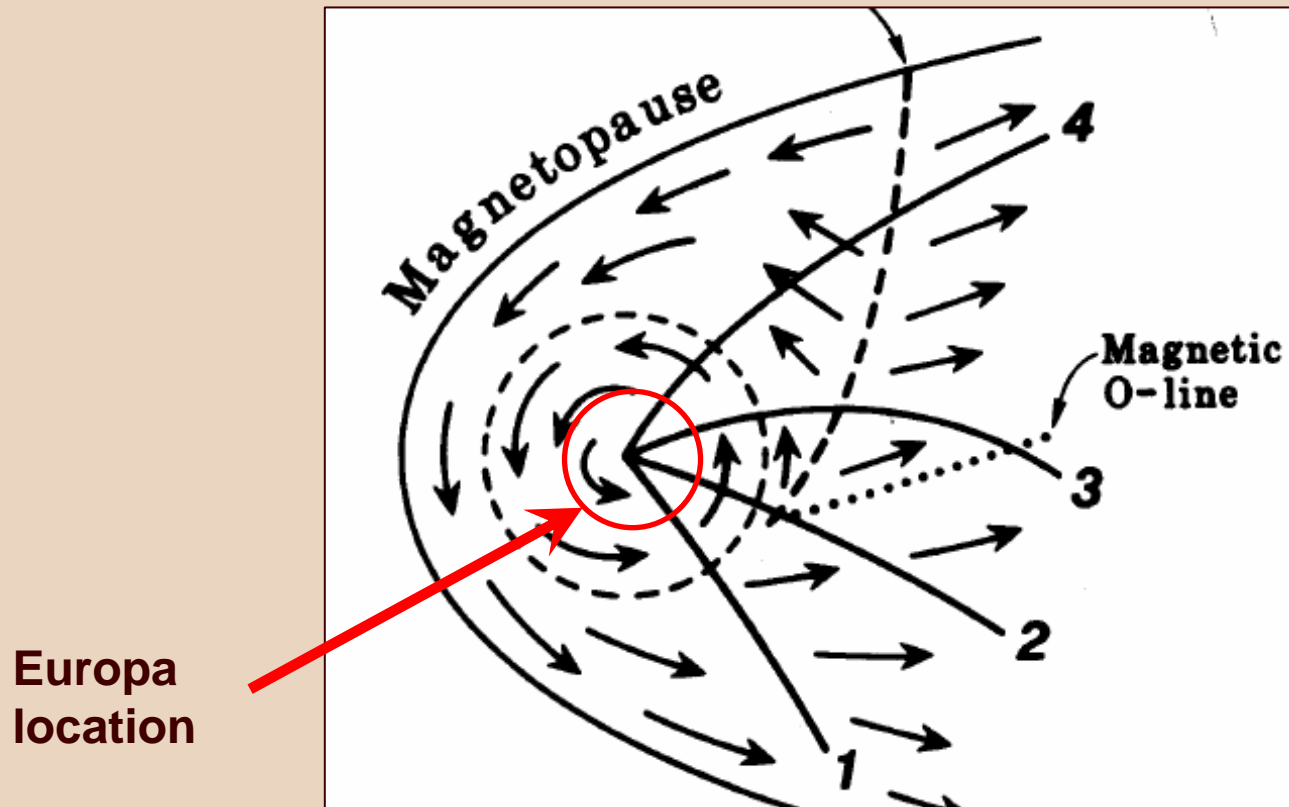
strong magnetic field & fast rotation: ★ powerful energizing environment

Io outflow: ★ strong outflow of cold heavy ions

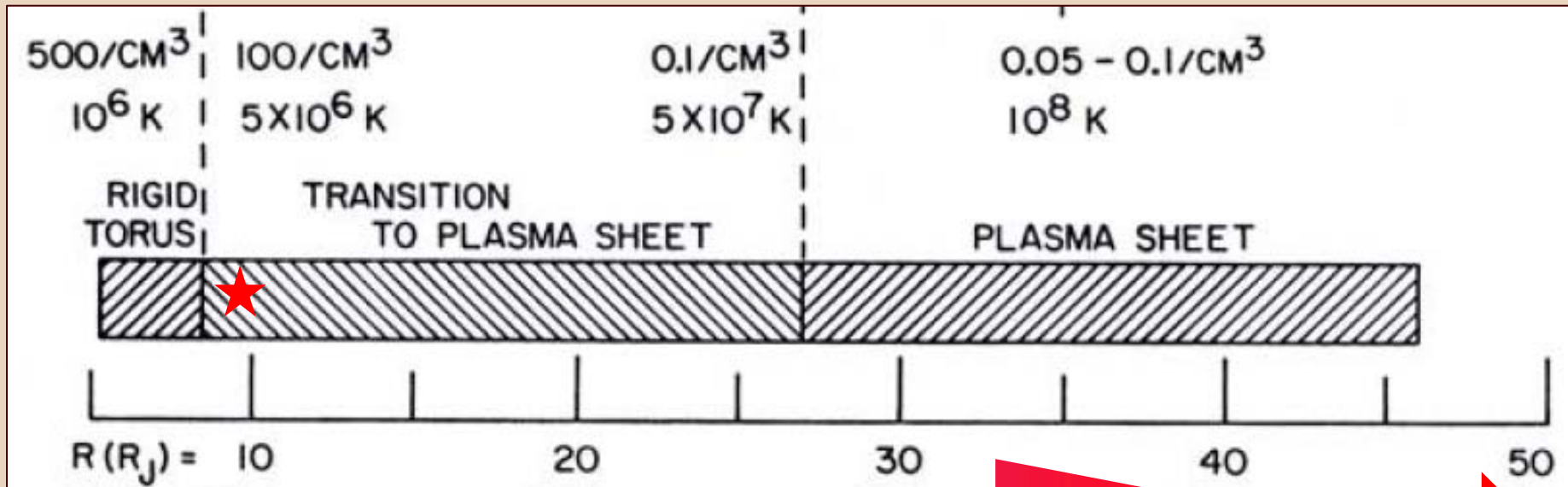
# rotating magnetosphere

electric potential of rotation at **Jupiter** is **400 MV**, solar wind “only” **1 MV**

at **Earth** both are  $\sim 100$  kV



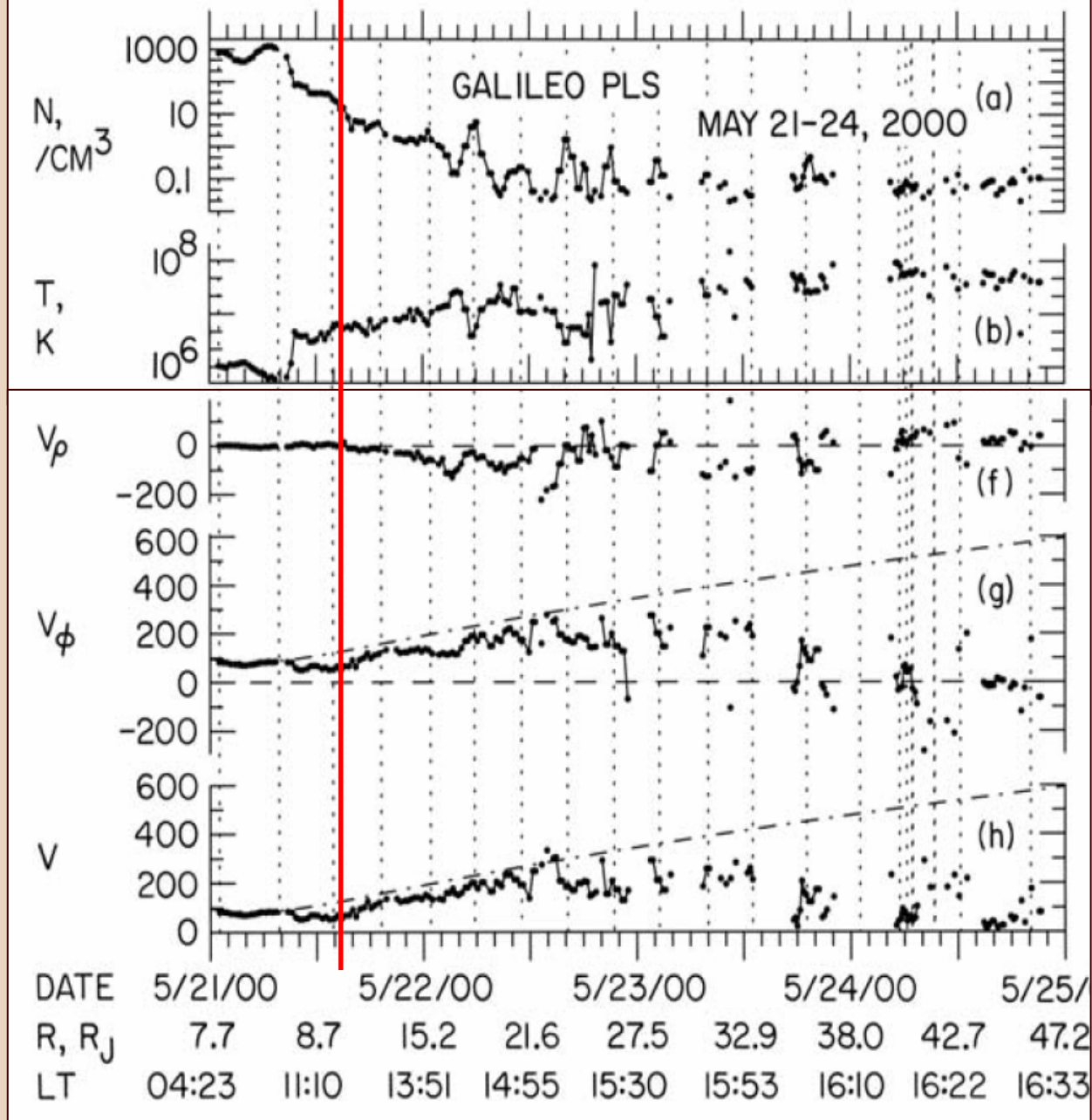
# particle populations



magnetic flux tubes with **cold plasma (100 eV)** diffuse out and heated  
most of plasma is lost

flux tubes with **hot rarefied plasma 10's of keV** return back to Jupiter

**Outer radiation belts ~ 1-100 MeV** are accelerated tails of hot plasma  
balance of hot and cold populations powers instabilities



# variability

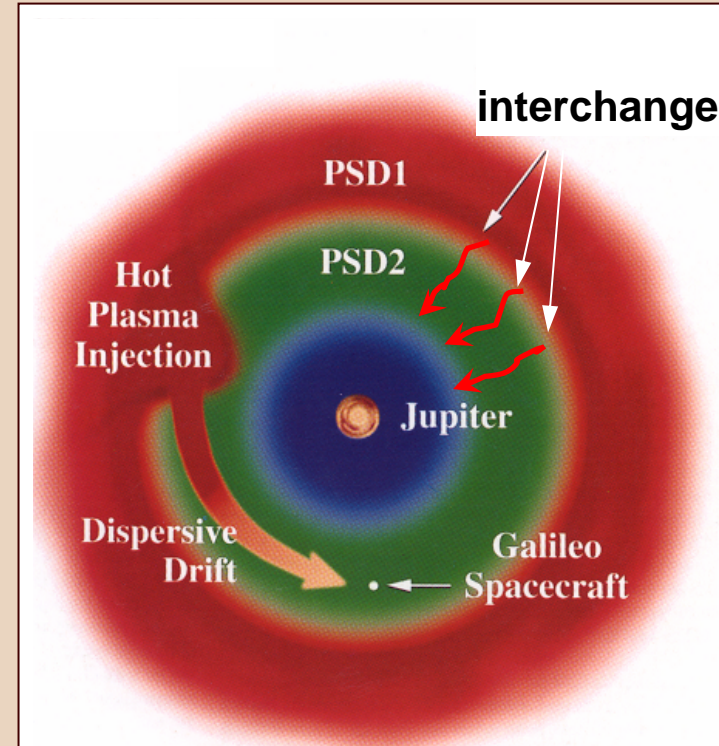
cold plasma diffuse outside + up and down motion relative to Europa orbit

bursty inward radial transport of hot plasma: three time scales

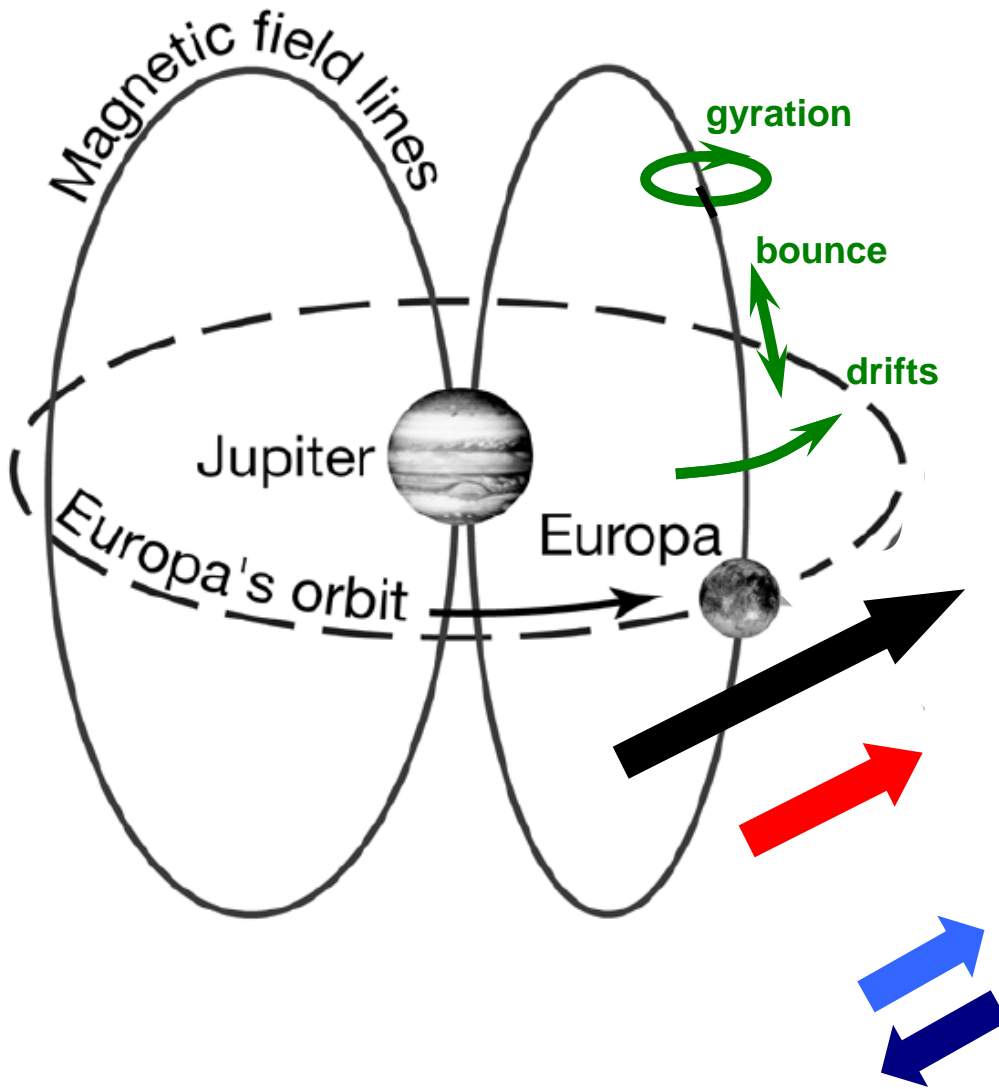
- interchange: relatively empty flux tubes penetrate inside  
**minutes, 100 km/s**
- injections: drifting hot plasma regions  
**hours**
- substorms: global unloading, a bunch of injections  
**days**

**cold plasma: 10-200 cm<sup>-3</sup> pressure 2-10 nPa**

**energetic plasma (> 20 keV): pressure ?-10 nPa**



# particle motion



Europa orbital motion 14 km/s

corotation velocity 117 km/s

**Europa is standing  
in plasma flow ~100 km/s**

**electric field drift ~ 40 mV/m**

**energetic plasma has gradient drifts**

**ions drift in direction of corotation**

Europa moves ~5000 km per bounce

electrons drift against corotation

**< 20 MeV can not overcome corotation**

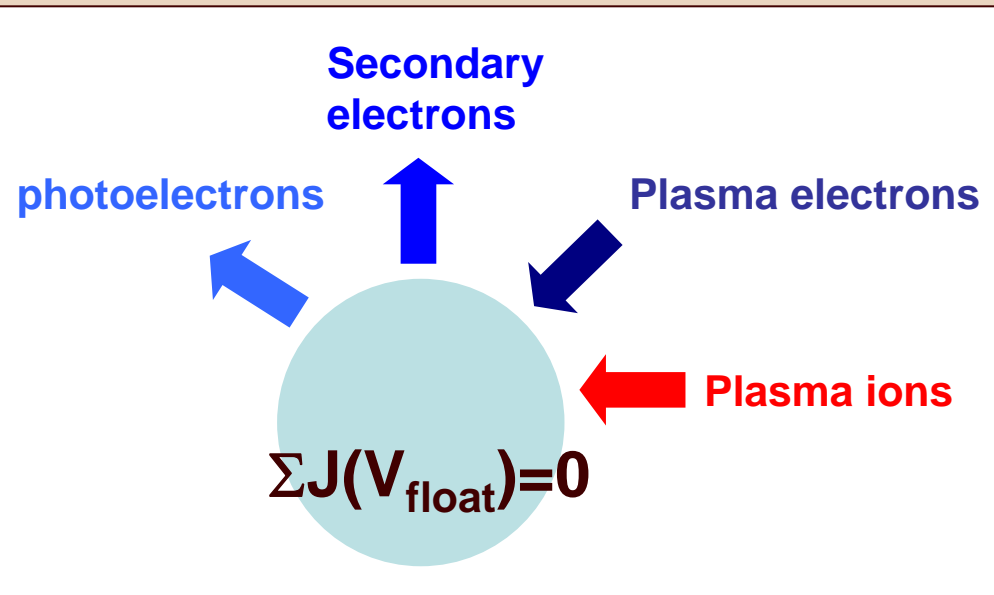
**> 20 MeV faster than corotation**

Europa moves ~300 km per bounce





## plasma effects-2



### ➤ surface charging

objects in plasma acquire floating potential so that total current is zero

floating potential is of the order of temperature of dominating flux

in solar wind (Phobos, asteroids) photoelectrons dominate  $V \sim +1-5 V$

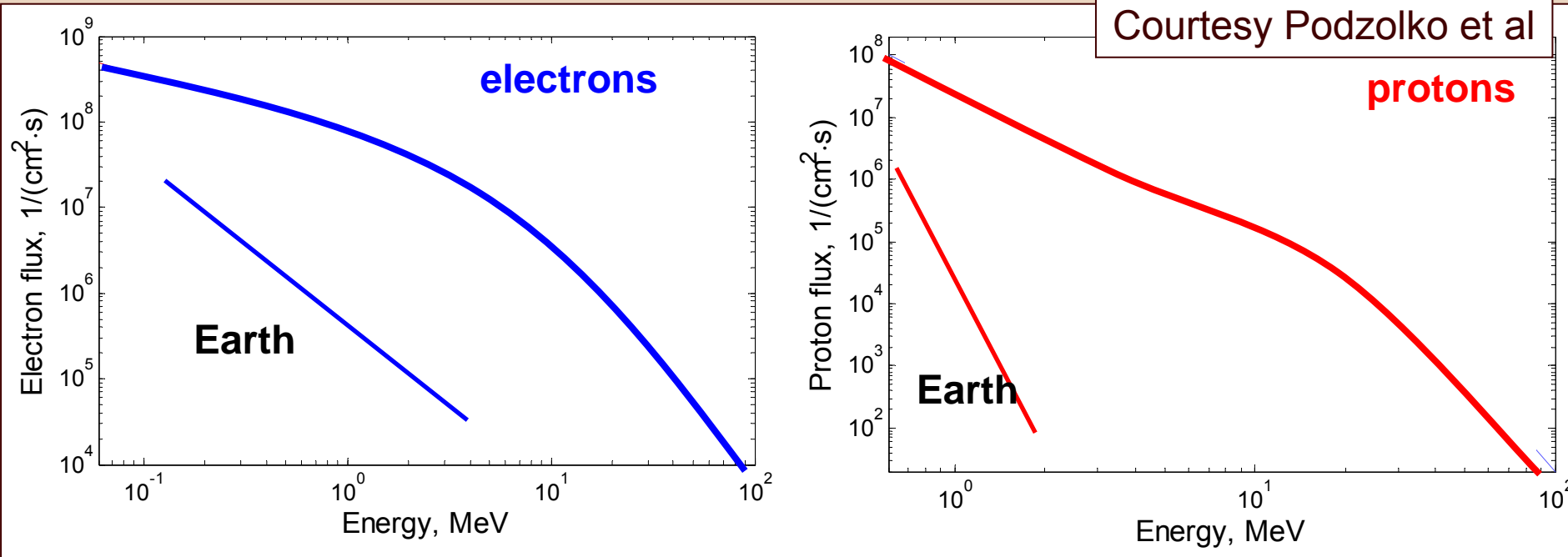
at Europa photocurrent is weak, if cold electrons dominate  $V \sim -100 V$   
if hot electrons (>keV)  $V \sim - \text{some keV}$

spacecraft are equipotential (by conducting layer)

but ice, regolith, moving elements may be differentially charged

# energetic plasmas

Courtesy Podzolko et al



total **dose** amounts to **megarad** - a killing quantity for a near-Earth satellites

**Galileo** survived 650 krad (estimate), 3 times above nominal,  
thanks to a very conservative design

for various instruments and conditions doses might be substantially different

dose is “global climate”, while “regional climate” and “weather” are also important

# plasma effects-3

dose **behind 1 cm** of Al is dominated by electrons: **1 Mrad** for 2 months

dose with **no shielding** is dominated by ions: **0.2 Grad** for 2 months

**ions** penetrate to sub-mm depths

important for open sensitive elements (like CCD's)

- ▶ total dose
- ▶ single event upsets

**electrons** pass ~0.5 cm of unit density per MeV

important for electronics

- ▶ total dose
- ▶ internal charging

**bremsstrahlung gamma rays** from electrons (>10 MeV) can pass 10's of cm



# plasma effects-4

- radiation affects the spacecraft as well as ... Europa surface

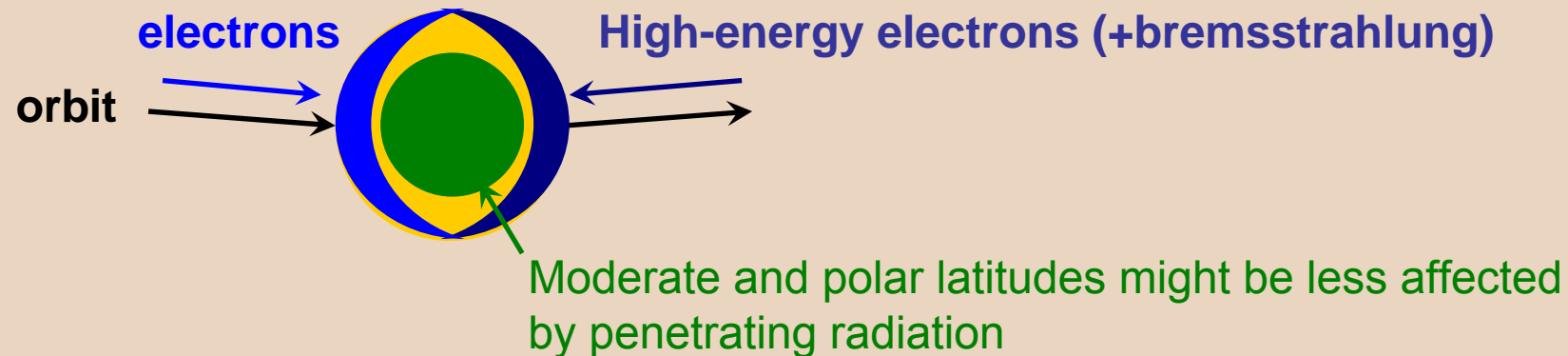
tomorrow's talk by Patterson et al.

1 rad = 100 erg/g ~  $10^{-10}$  eV/nuclon

\* mega/gigarads \* millions of years

what organics can survive ? How deep to dig ?

detailed calculations + surface formation + differential rotation



# conclusions

Jupiter has powerful magnetosphere

- ✓ a lot of heavy ions
- ✓ fast acceleration of new ions
- ✓ surface charging
- ✓ radiation effects on spacecraft
- ✓ radiation effects on surface

