

Space Weather

— History and Current Status

Ji Wu

National Space Science Center, CAS

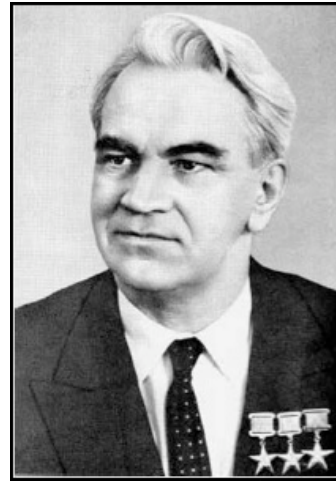
Oct. 3, 2017

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1. Beginning of Space Age and Dangerous Environment

Space Age



Kai'erdishi



Korolev

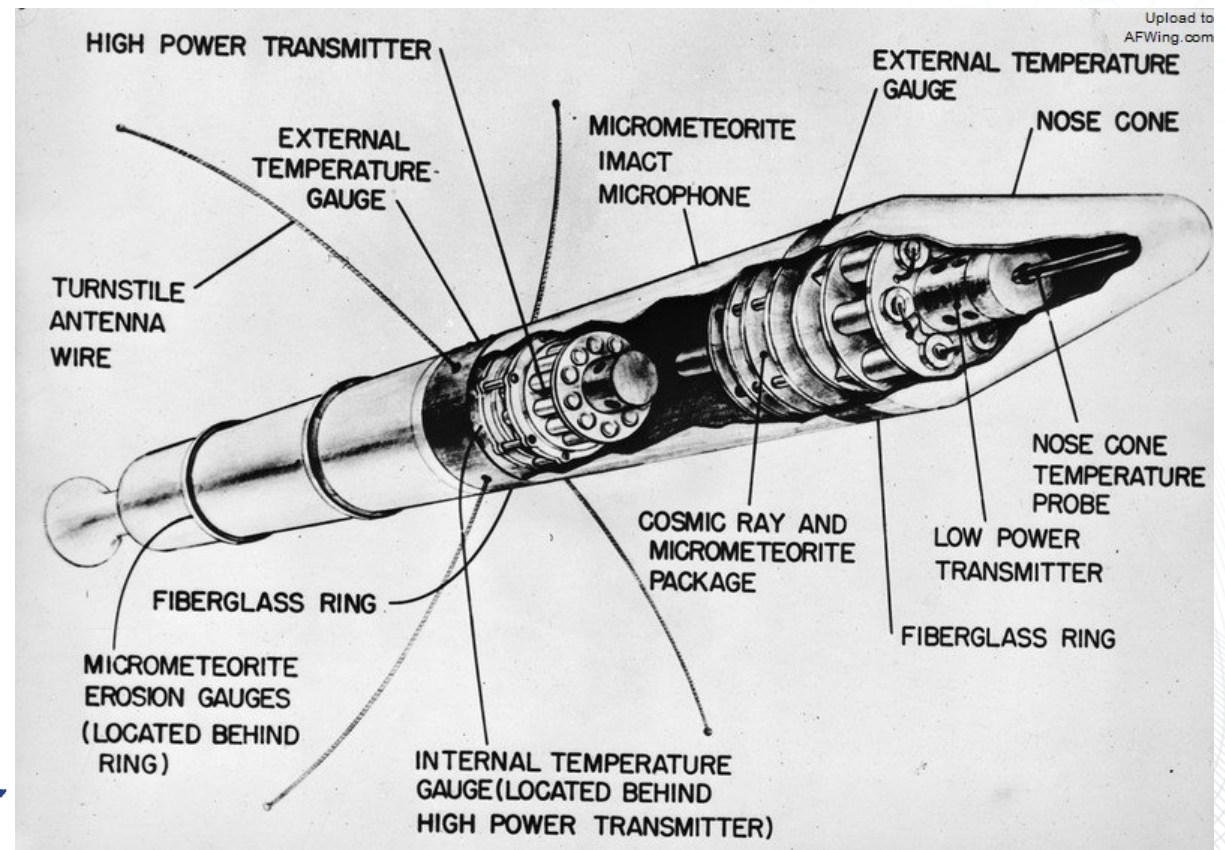


Oct. 4, 1957, humanity's first artificial satellite, **Sputnik-1**, has launched, ushering in the Space Age.

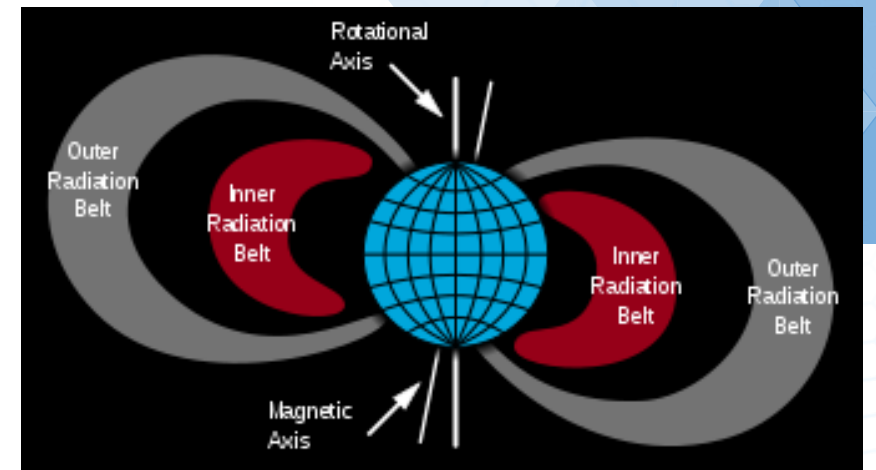
Space Age



Explorer 1 was the first satellite of the United States, launched on Jan 31, 1958, with scientific object to explore the radiation environment of geospace.

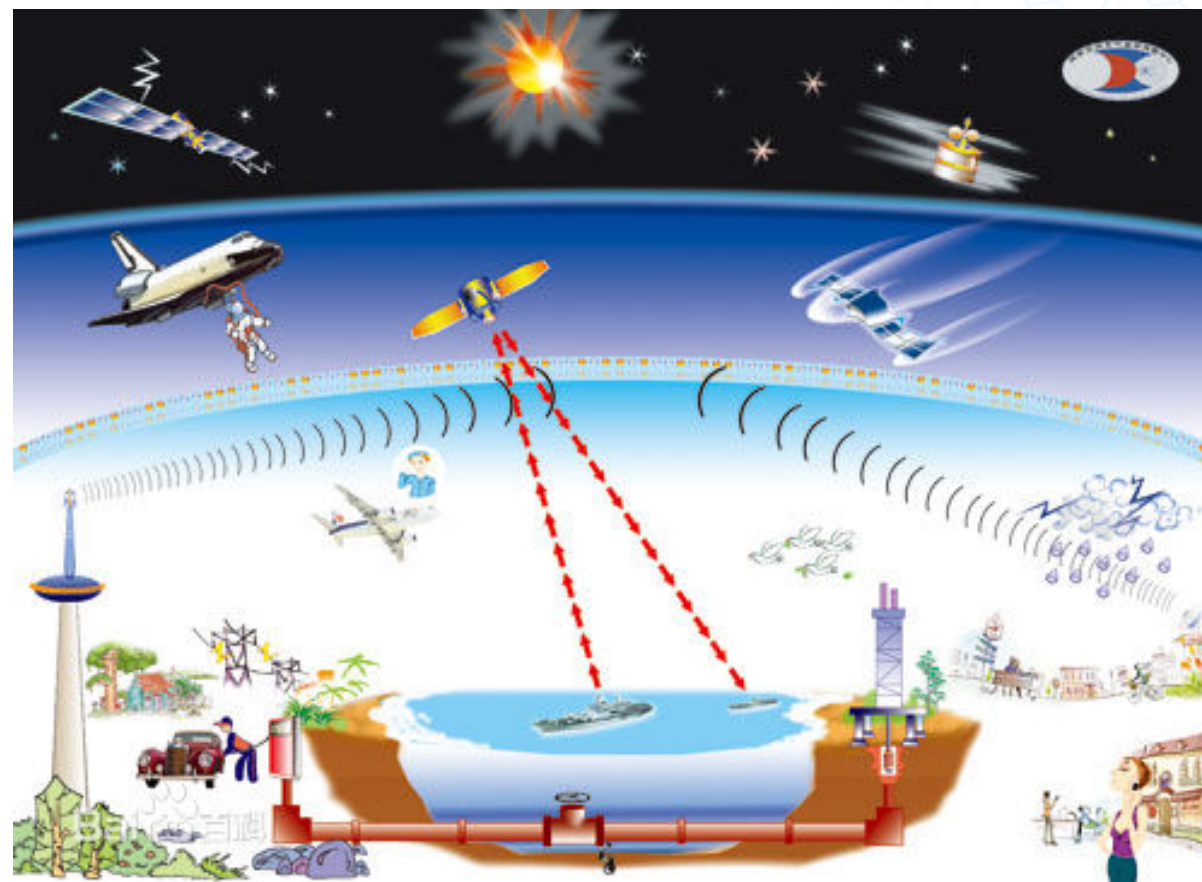
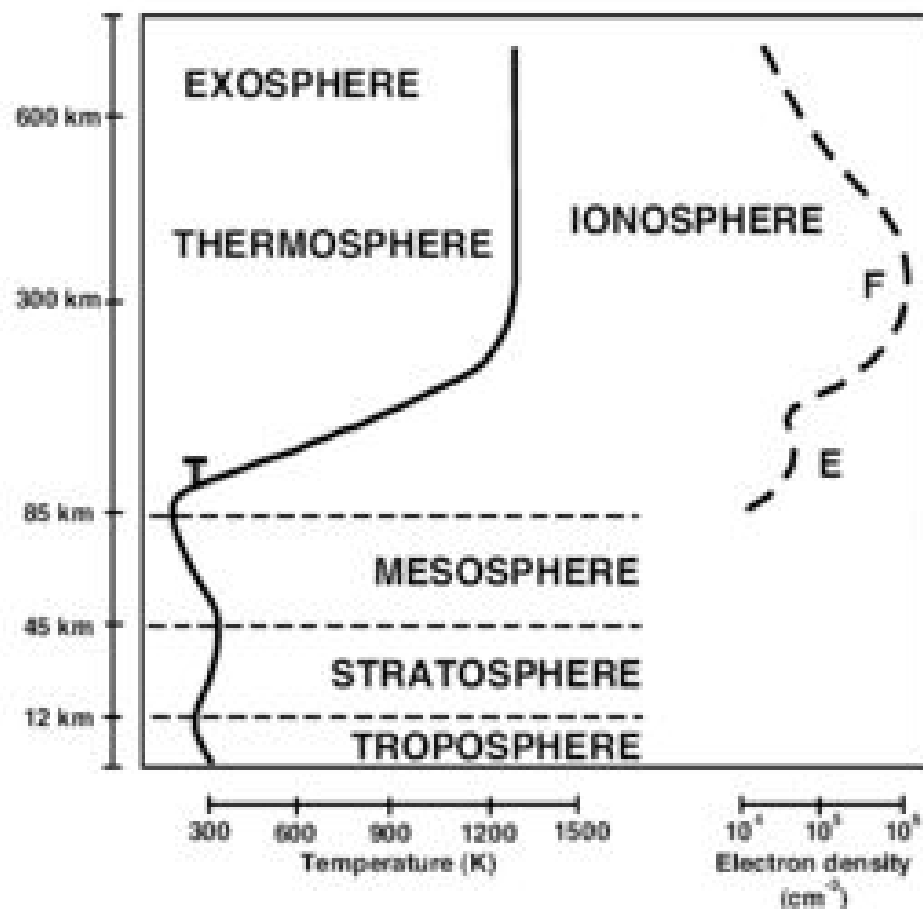


Unknown Space Environment



- **Sputnik-2** (Nov 3, 1957) detected the Earth's outer radiation belt in the far northern latitudes, but researchers did not immediately realize the significance of the elevated radiation because Sputnik 2 passed through the **Van Allen belt** too far out of range of the Soviet tracking stations.
- **Explorer-1** detected fewer cosmic rays in its orbit (which ranged from 220 miles from Earth to 1,563 miles) than Van Allen expected.

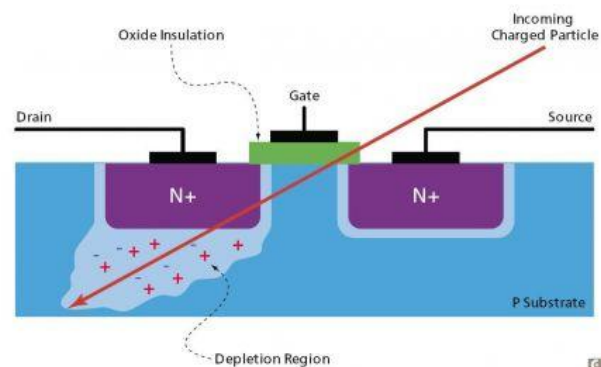
Space Age - unknown and dangerous space environment



Satellite failures due to the unknown and dangerous space environment

Particle Radiation !

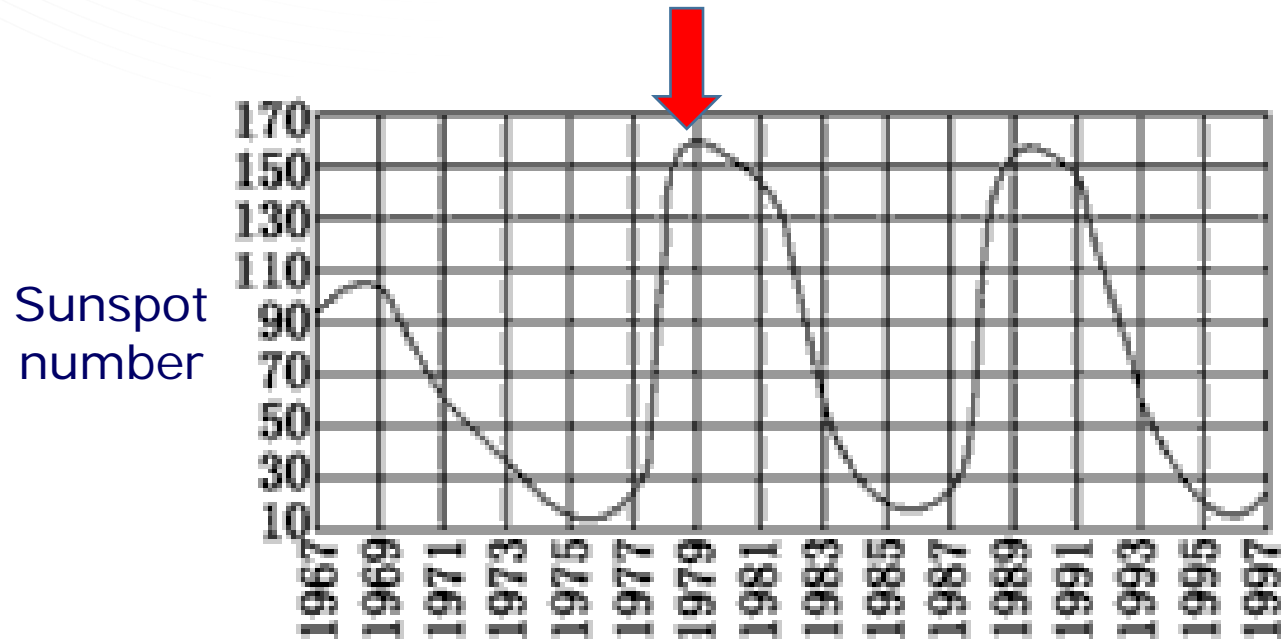
- Statistics show that the space radiation environment is one of the **main causes of satellite failure**.
- The space radiation environment caused about **2,300** satellite failures of all the **5000** failure events during the 1966-1994 period collected by the National Geophysical Data Center.
- Statistics of the United States in 1996 indicate that the space environment caused more than **40%** of satellite failures in 1958-1986, and **36%** in 1986-1996.



Space Environment speeded up the falling down of the Skylab

Magnetic Field and Atmosphere !

In 1979, the Skylab space station succumbed to the long-term effects of atmosphere drag and plunged back to earth.



Environment can be very serious!

1989 Geomagnetic Storm

- A severe geomagnetic storm struck Earth on March 13, 1989. It occurred during solar cycle 22 and caused a nine-hour outage of Hydro-Québec's electricity transmission system.
- This storm bringing down the Galaxy 4 satellite, halting news transmissions and electronic pagers across North America for days.



Photograph of the aurora of March 13, 1989, taken from, NY



Thousands of users lost contact from Galaxy 4 Comm. Sat.

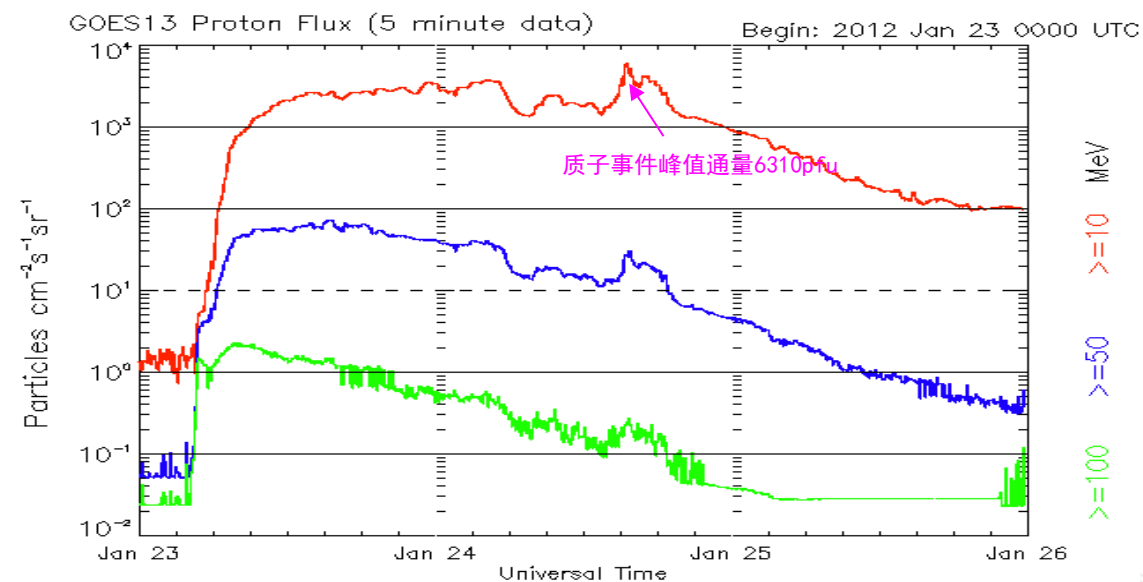
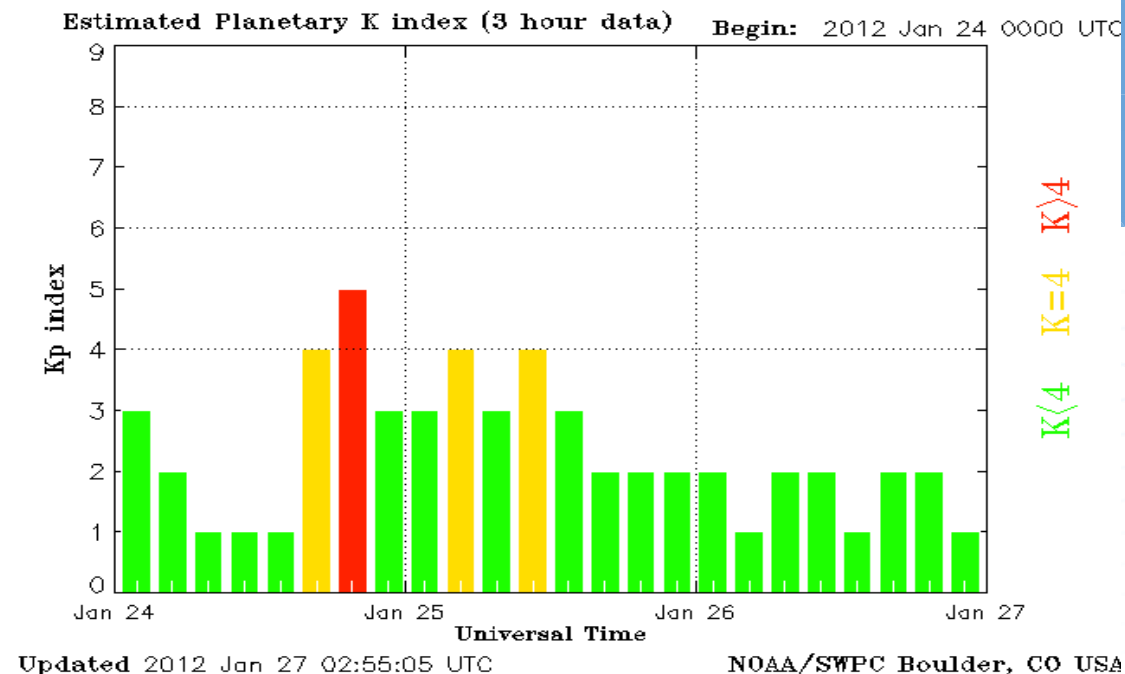
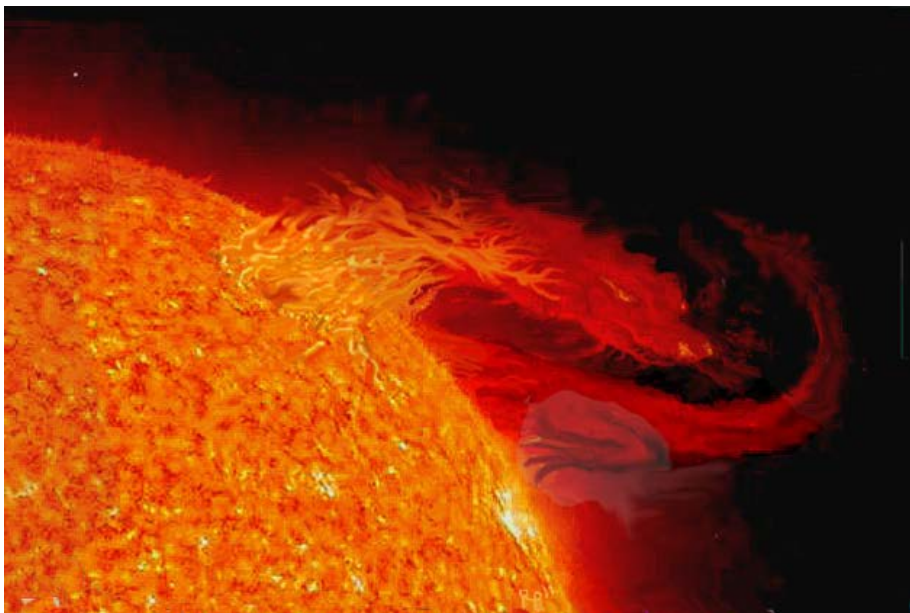


PJM Public Service
Step Up Transformer
Severe internal damage caused by
the space storm of 13 March, 1989



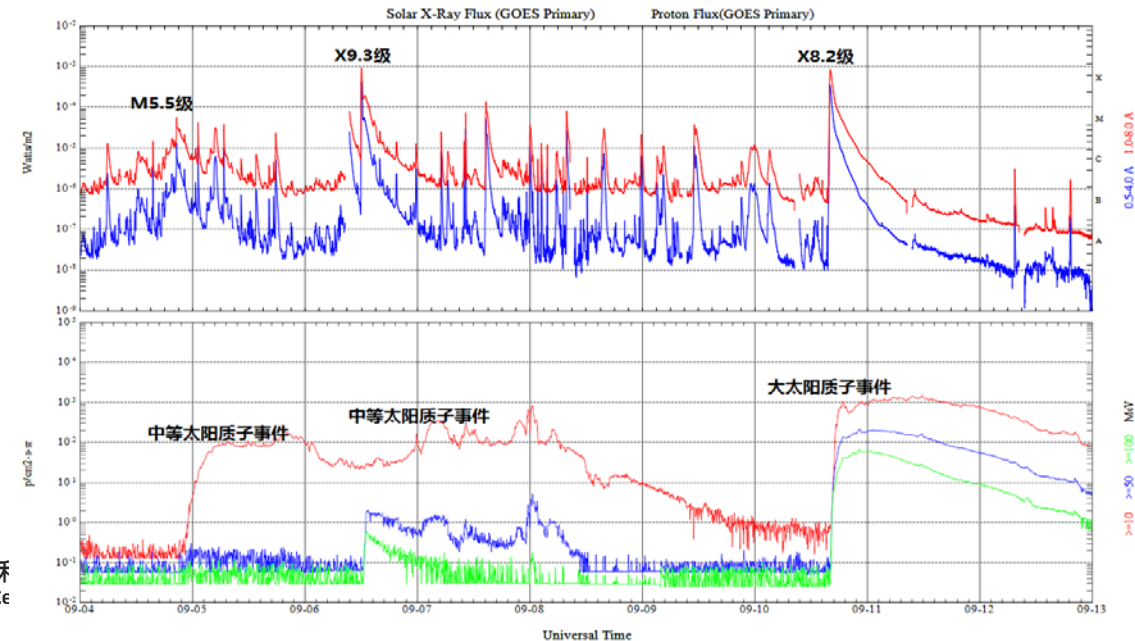
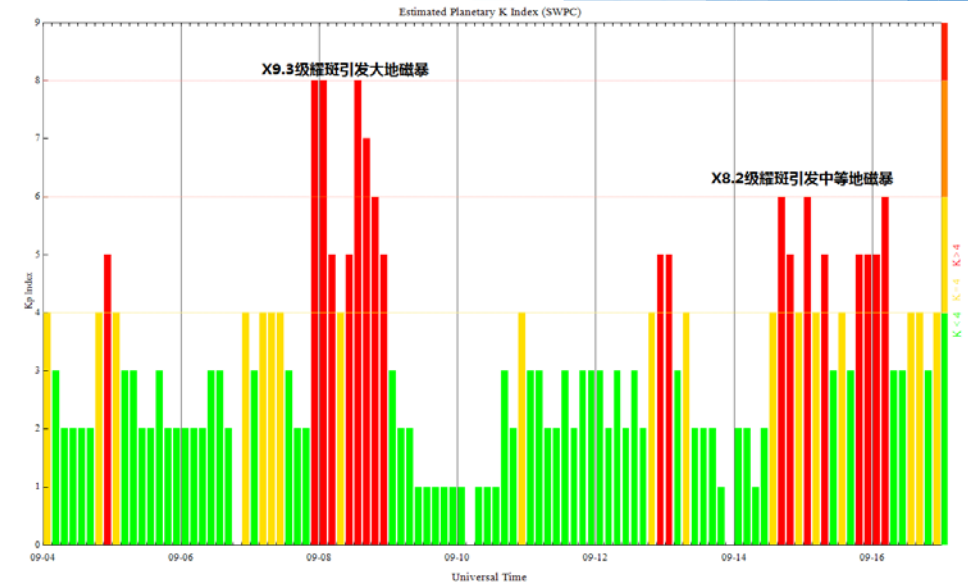
China Dragon Event

- AR1402 bursted a M8.7 flare with solar proton event on **Jan 23 in 2012**. The flux of solar proton event reached 6310pfu, high-speed CME reached earth in 1.5 days and caused the geomagnetic disturbance.



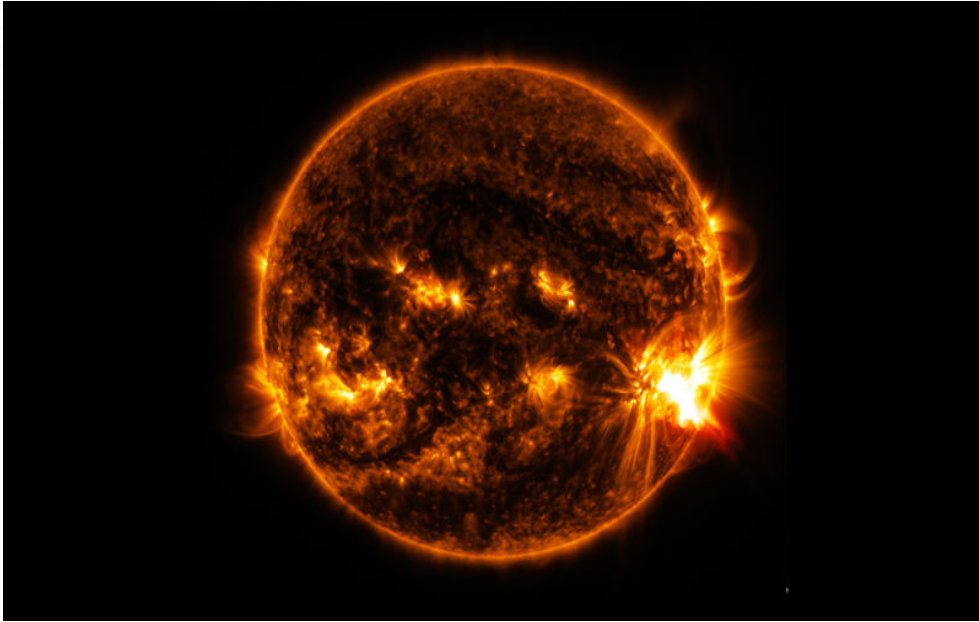
Hungry Ghost Festival Event

- At 7:53 in the evening of September 6, 2017, a large flare (X9.3) triggered solar proton events and CME. It was the strongest solar activity since 2005, and fired the first shot of a new solar storm.
- The second day of this event coincides with the traditional festival – ‘Hungry Ghost Festival’, so the name of this great event was named as ‘Hungry Ghost Festival Event’.



The important fact is:

We have been in space for only **60** years, nevertheless the Sun was there for several billions years already!



We never know if we will get a much stronger Solar Storm tomorrow!

Artist's concept depicting energetic particles from solar super flares raining down on the early Earth.

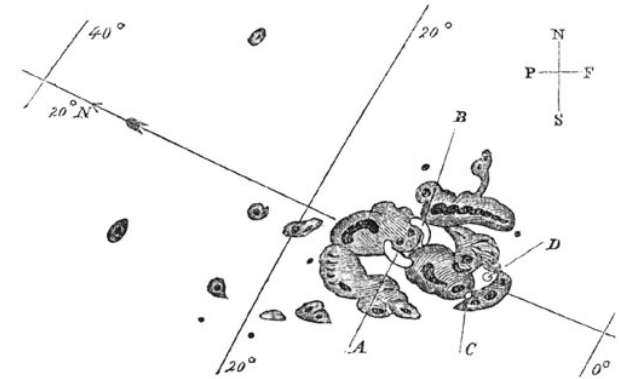


Image via Vladimir Airapetian

The Carrington Event

A solar coronal mass ejection hit Earth's magnetosphere and induced one of the largest geomagnetic storms on record, September 1–2, 1859.

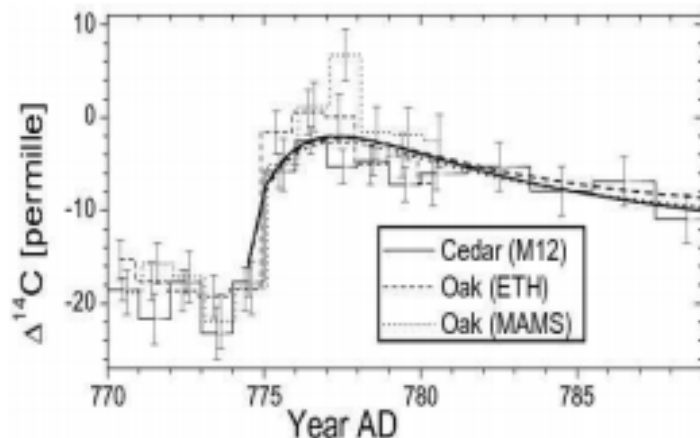
- Telegraph systems all over Europe and North America failed;
- Southern auroras were observed as far north as Queensland, Australia



Sunspots of September 1, 1859, as sketched by Richard Carrington.

AD775 Event

In the evening on the Chinese lunar calendar day of 11 Dec. 774, i.e., 17 Jan. AD775, in the east and above Moon, there were more than ten bands of white lights like the spread silk, penetrating and covering eight grand constellations named in Chinese. The lights were ceased gradually after middle night, as recorded in the Old Tang Book - a Chinese.



33RD INTERNATIONAL COSMIC RAY CONFERENCE, RIO DE JANEIRO 2013
THE ASTROPARTICLE PHYSICS CONFERENCE

ICRC
2013

The Solar Cosmic-Ray Origin for the Rapid ^{14}C Increase in AD775

D. ZHOU¹, C. WANG¹, Z. PENG², R. RUTLEDGE³, Y. SUN¹, J. LIANG¹, G. ZHU¹, S. ZHANG¹, B. ZHANG¹, P. ZHOU¹, J. WU¹

¹ National Space Science Center, Chinese Academy of Sciences, Beijing 100190, China

² University of Science and Technology of China, Hefei 230026, China

³ NOAA - Space Weather Prediction Center, Boulder, CO 80505, USA

dazhuangzhou@gmail.com

Abstract: The rapid ^{14}C found. The origin of ^{14}C event. There was no detected the energy needed for light years away and the be detected, however in the ^{14}C increase in AD775. CMEs (Coronal Mass Ejections) generated by the recorded in the Old Tang Book indicated the physical particles. Theoretical models which explode towards Therefore, the big ^{14}C emission, or in another

Keywords: cosmic ray

Chin. Sci. Bull. (2014) 59(22):2736–2742
DOI 10.1007/s11434-014-0345-2

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www.springer.com/csb

High-Energy Physics

Super solar particle event around AD775 was found

Dazhuang Zhou · Chi Wang · Binqun Zhang ·
Shenxi Zhang · Ping Zhou · Yueqiang Sun ·
Jinbao Liang · Guangwu Zhu · Ji Wu

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Abstract A rapid radiocarbon ^{14}C increase of 12‰ in AD774–775 has been reported in cedar and oak tree rings. So far, the origin of the ^{14}C increase is still uncertain and the possible origin is either supernova or solar particle event. The most possible origin of ^{14}C increase is strong solar flares and Coronal Mass Ejections (CMEs) with strong particles emission. Comprehensive approaches to identify the strong historical solar particle events based on the rapid ^{14}C increase in tree and coral rings and ice cores, long duration strong auroras and geomagnetic storms are introduced. Evidence of the super auroras in AD775 was first found in a Chinese Chronicles Jiutangshu and it supports the views that the rapid ^{14}C increase and strong auroras around AD775 are most possibly caused by strong solar storms with intense particles emission. It was identified that the solar event around AD775 would be the strongest solar particle event in the past 11400 years. The discovery is significant for the research on the history of solar activities, space weather and forecast, radiation of solar energetic particles and protection.

Keywords Cosmic rays · Solar events · Radiocarbon · Paleoclimatology · Space weather

1 Introduction

Galactic cosmic rays (GCRs) are originated from the space out of solar system and solar energetic particles (SEPs) are

D. Zhou (✉) · C. Wang · B. Zhang · S. Zhang · P. Zhou · Y. Sun · J. Liang · G. Zhu · J. Wu
National Space Science Center, Chinese Academy of Sciences,
Beijing 100190, China
e-mail: dazhuangzhou@gmail.com

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originated from high energy solar activities—solar flares and CMEs. GCRs and SEPs are composed of protons, heavy nuclei and electrons. The ^{14}C on Earth can be produced by interactions between GCRs/SEPs and the Earth's atmosphere, and auroras can be produced by interactions between SEPs and the Earth's atmosphere and geomagnetic fields.

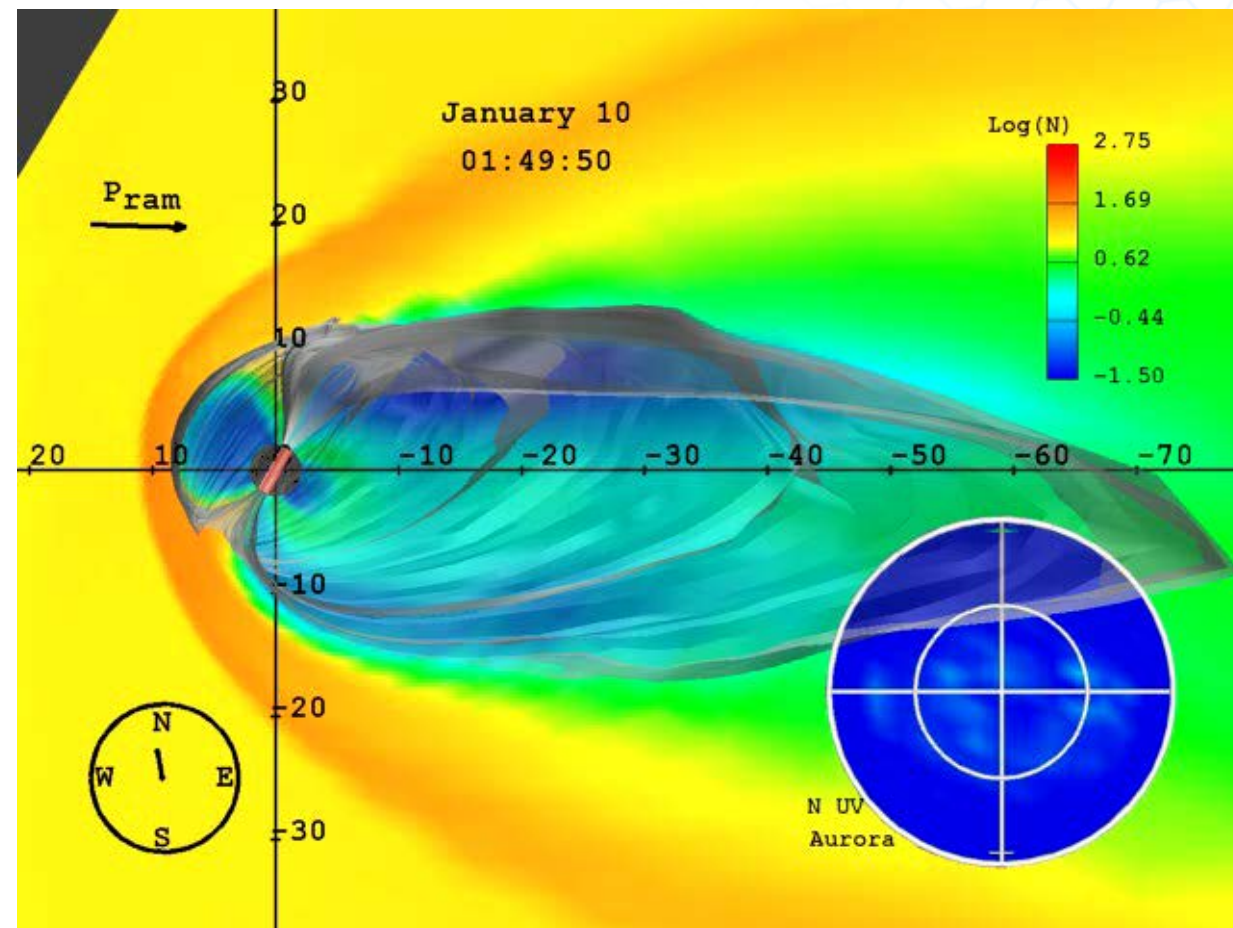
Recent research indicates that there was a rapid ^{14}C increase in tree and coral rings around AD775 [1–3]. The ^{14}C increase cannot be explained by normal GCRs and supernova. It is most possibly generated by strong solar event with intense particles emission. Although, there have been many experimental and theoretical research results related to the ^{14}C increase around AD775, so far the origin of ^{14}C is still uncertain and the cause — solar particle event, is still only an assumption or an analytical result due to the lack of evidence for auroras around AD775 [1–5]. The AD775 solar particle event is the candidate for the strongest solar particle event in the past 11400 years according to the ^{14}C measurements [2]. In fact, the important evidence—the world-wide auroras around AD775 can be found and identified with the historical records accumulated in the world, especially in China [6].

The information of historical solar particle events is significant in research on the historical solar activities, generation and acceleration of solar particles, space weather and forecast, and radiation of SEPs and protection.

This paper summarizes briefly the experimental results of the rapid ^{14}C increase around AD775, introduces the comprehensive approaches to identify strong historical solar particle events and the approach to identify the origin of ^{14}C increase and auroras, presents the new evidence of auroras around AD775 and the discovery of AD775 strong solar particle event as well as some new results related, including the fatal radiation of AD775 solar particles.

Zhou, et al, 2013

2. The Dynamic Space Environment So Far as We Know



Solar-Terrestrial Physics

■ Sun

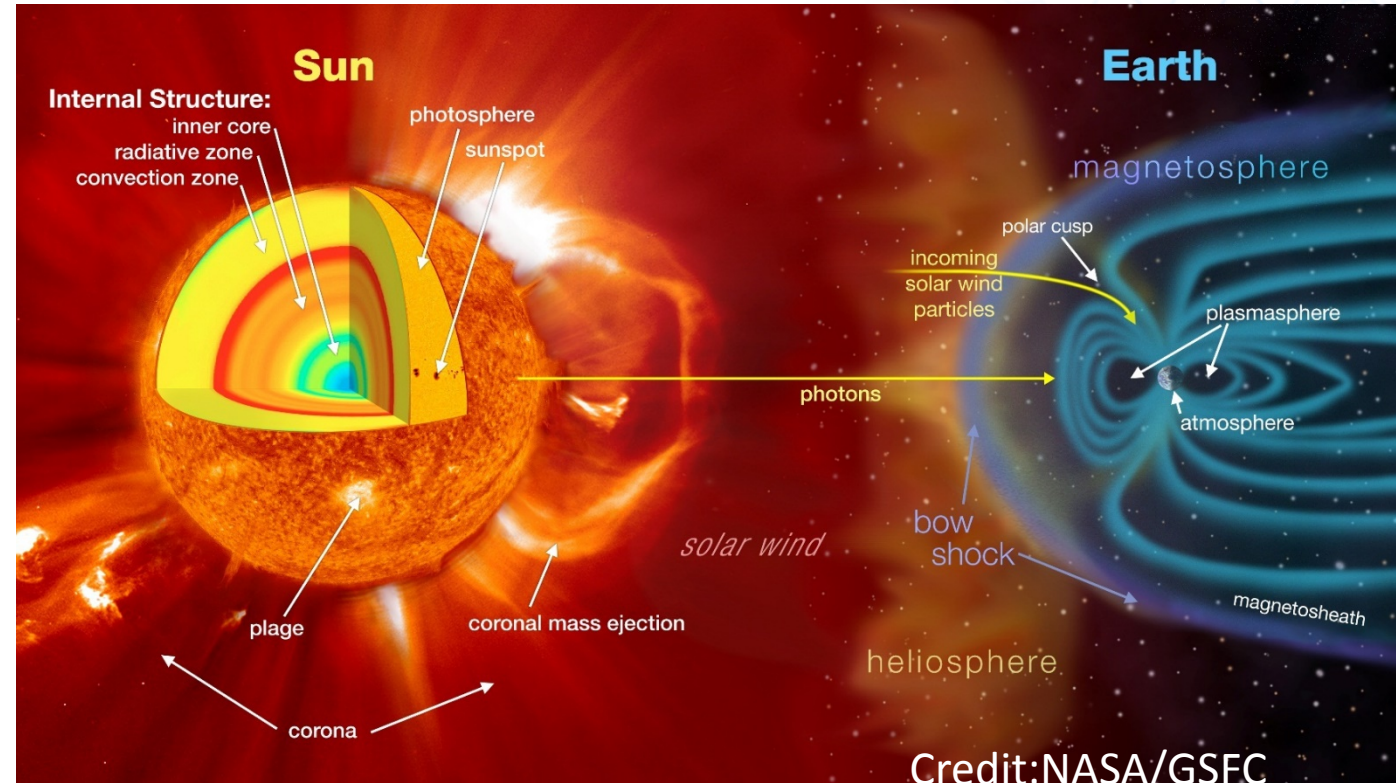
■ Solar Wind/Interplanetary

■ Geospace Environment

–Magnetosphere

–Ionosphere

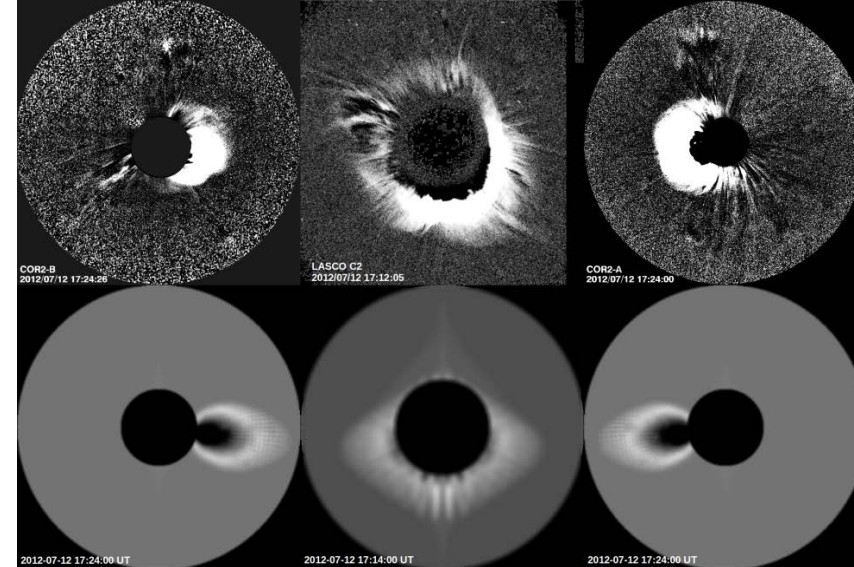
–Thermosphere



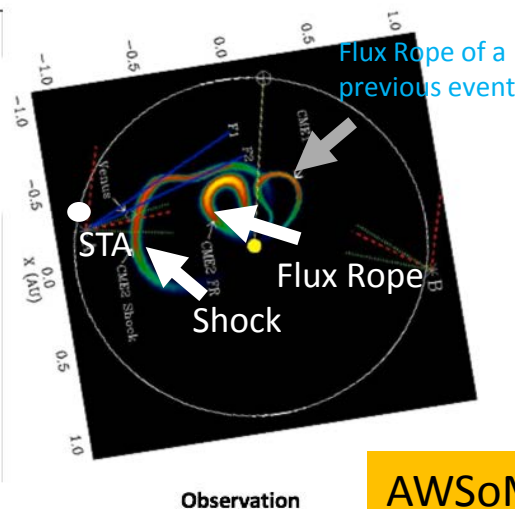
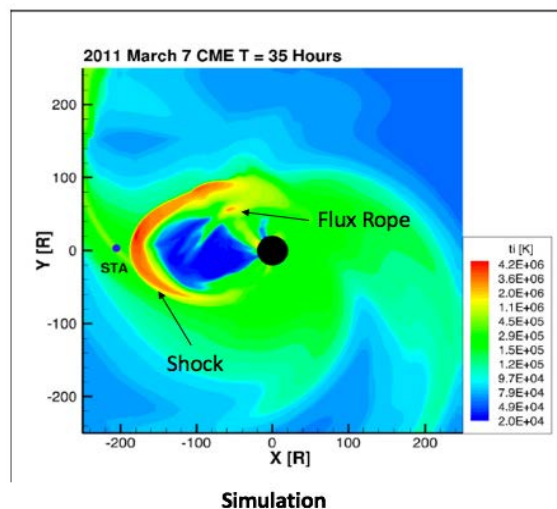


Coronal Mass Ejection

Significant progress has been made in numerical simulation of CME events in recent years.

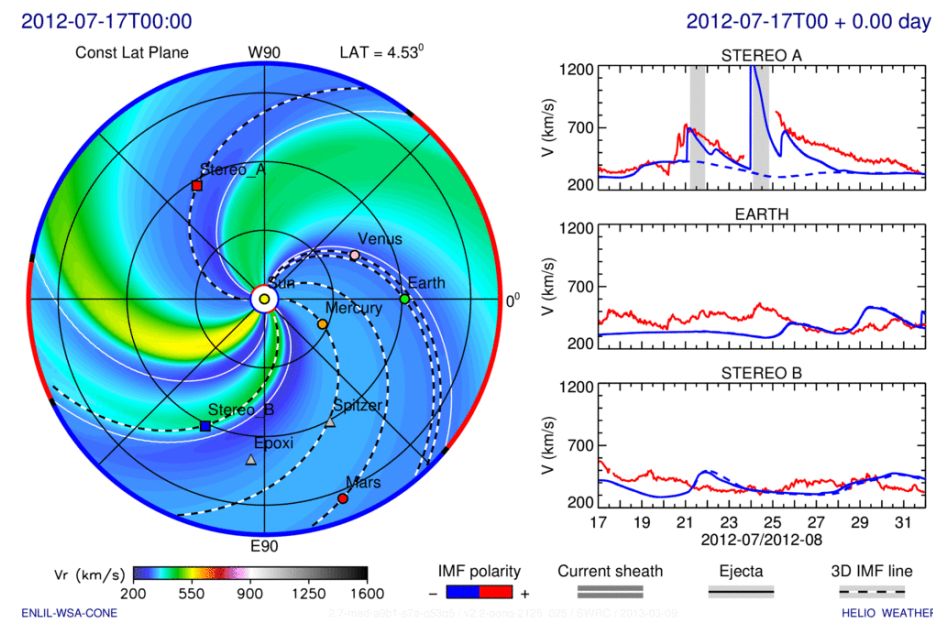


COIN model(SIGMA Group, NSSC)



AWSOM

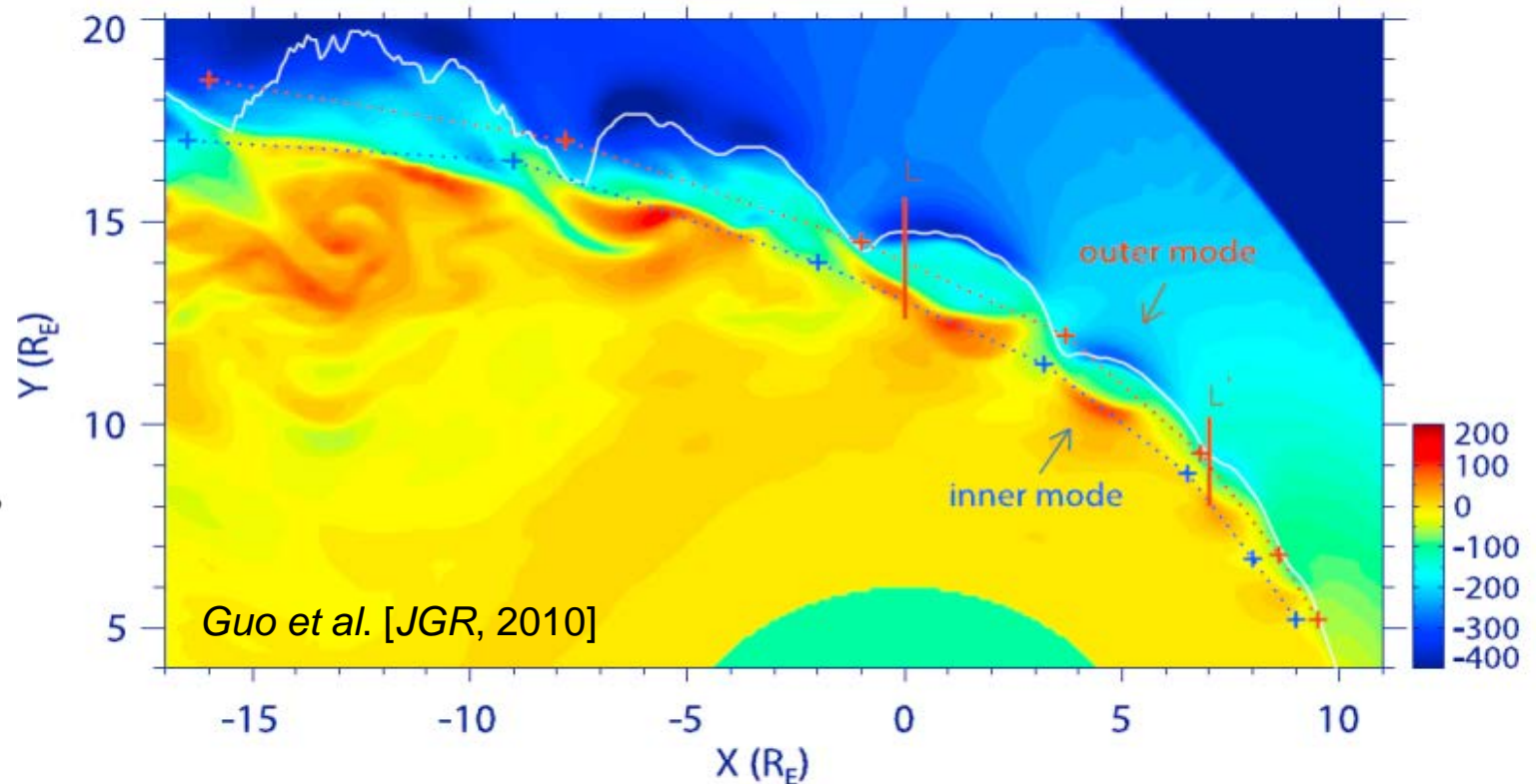
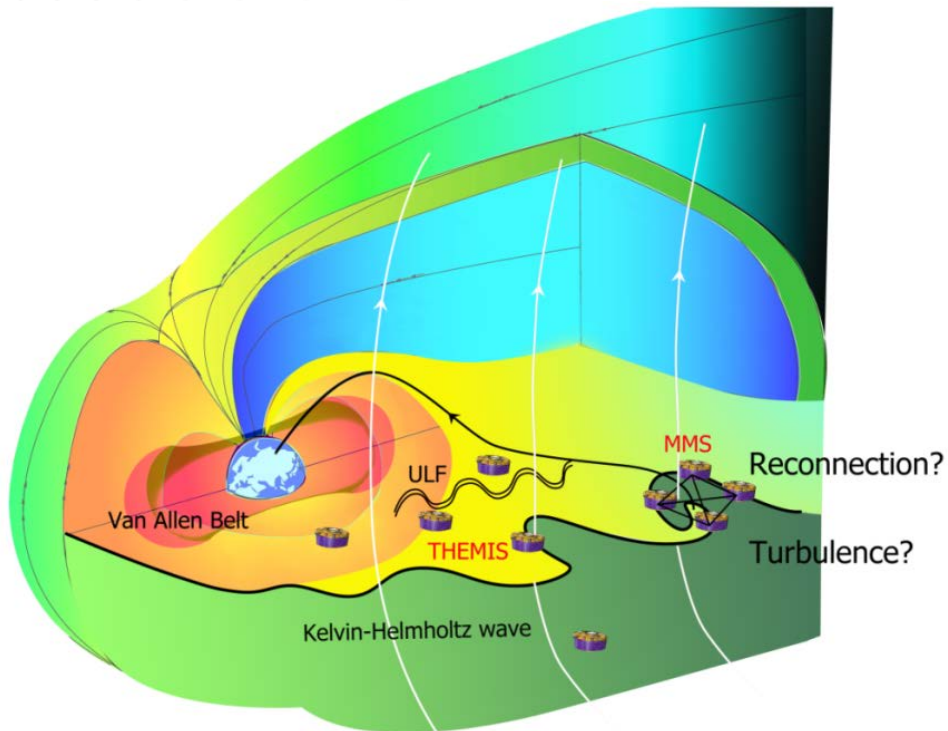
AWSOM model(Michigan University)



ENLIL model(US)

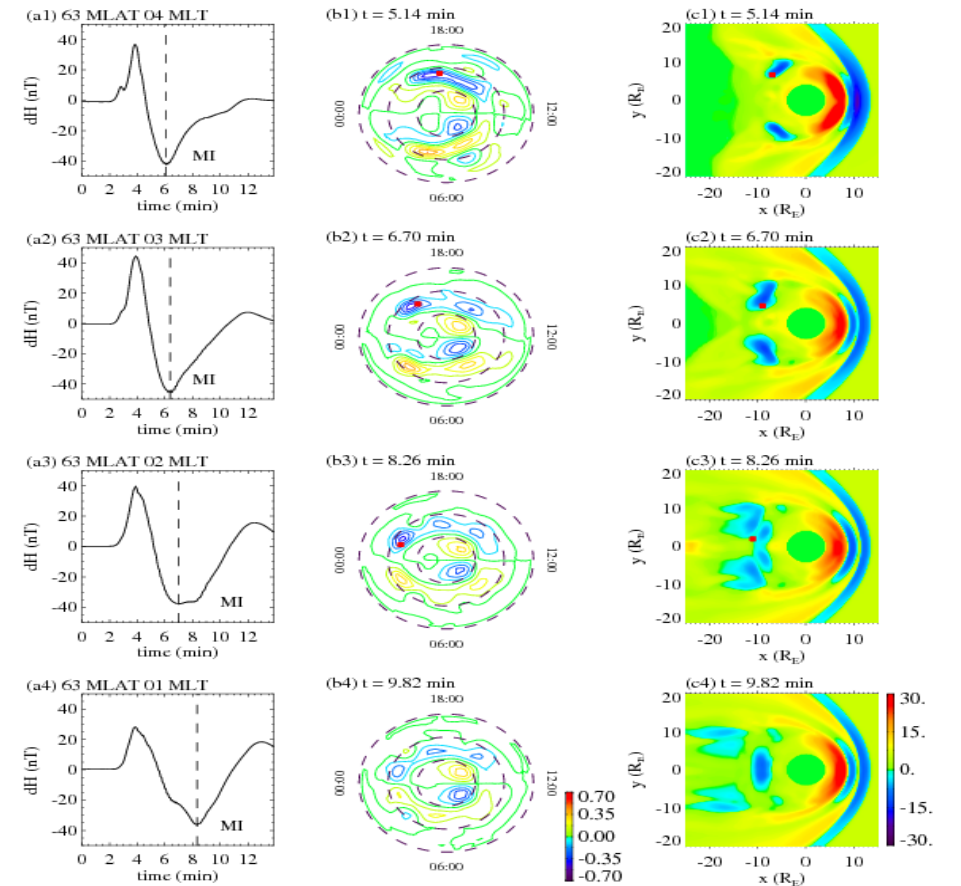
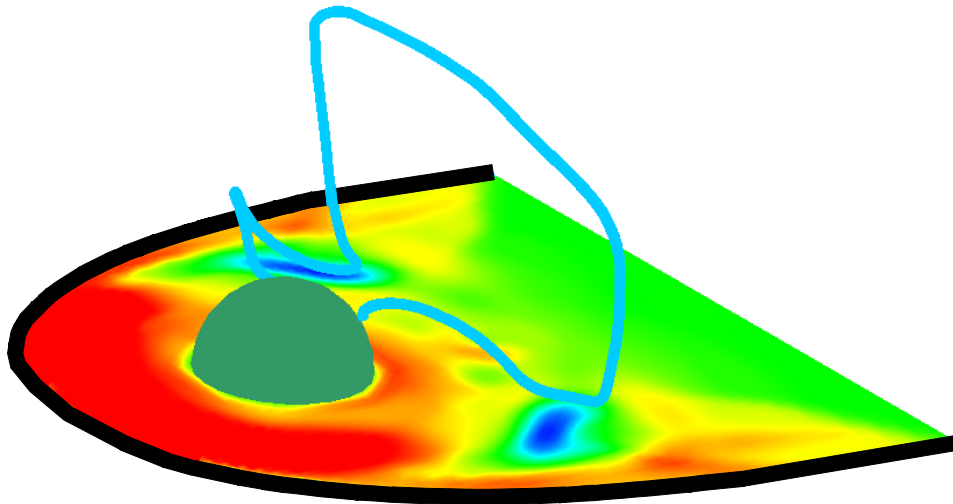
Magnetopause Kelvin-Helmholtz Instability

Deepen the understanding of the magnetopause instability



Magnetosphere Respond to Interplanetary Disturbance

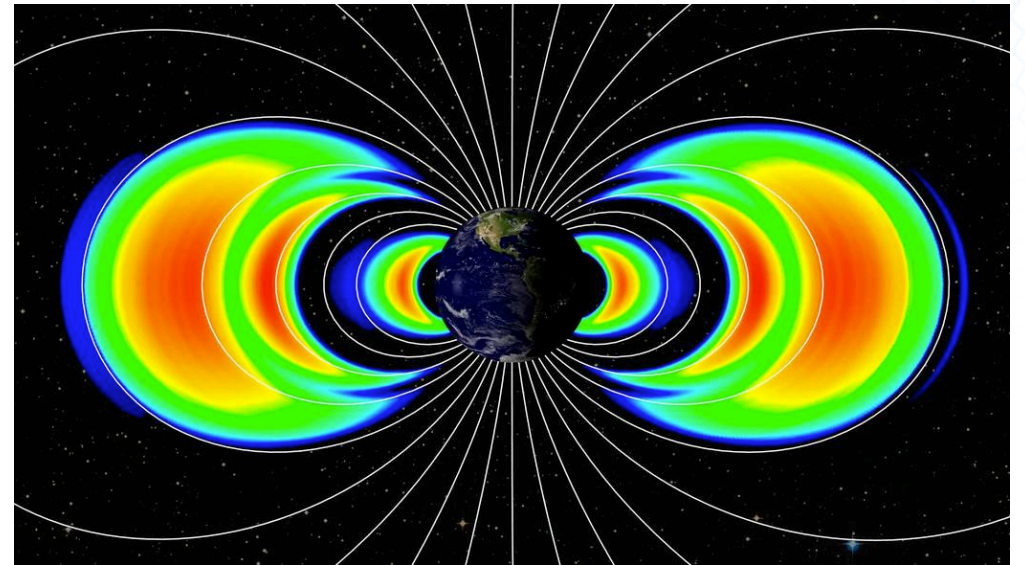
Solar Wind-Magnetosphere- Magnetic Field on Earth response link.



Sun et al., JGR, 2015

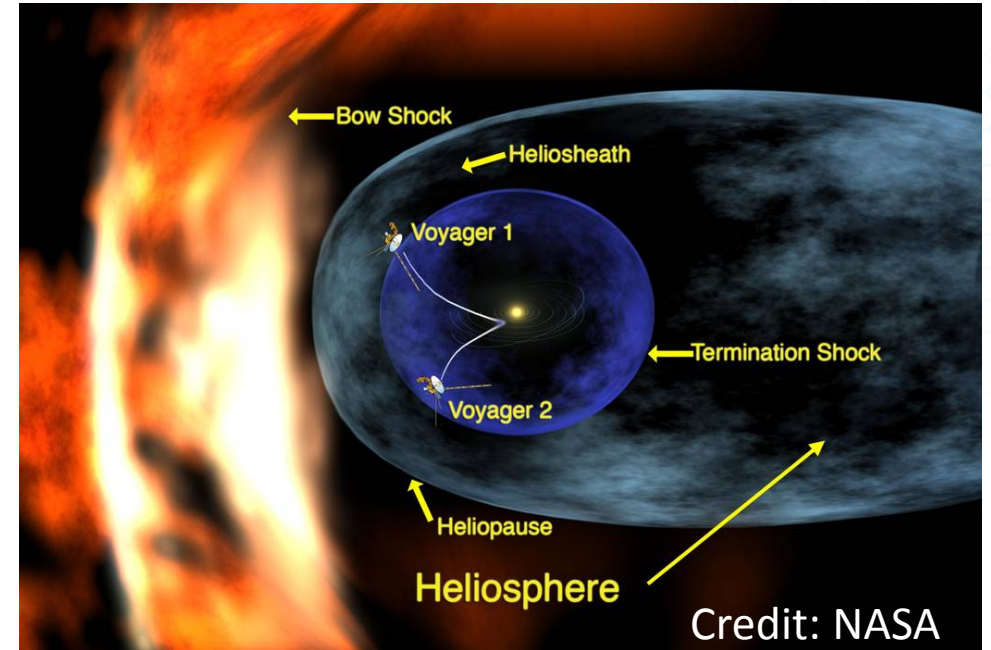
'NEW' Van Allen Belt

2013, A new radiation belt has been discovered around Earth by the Van Allen Probes.



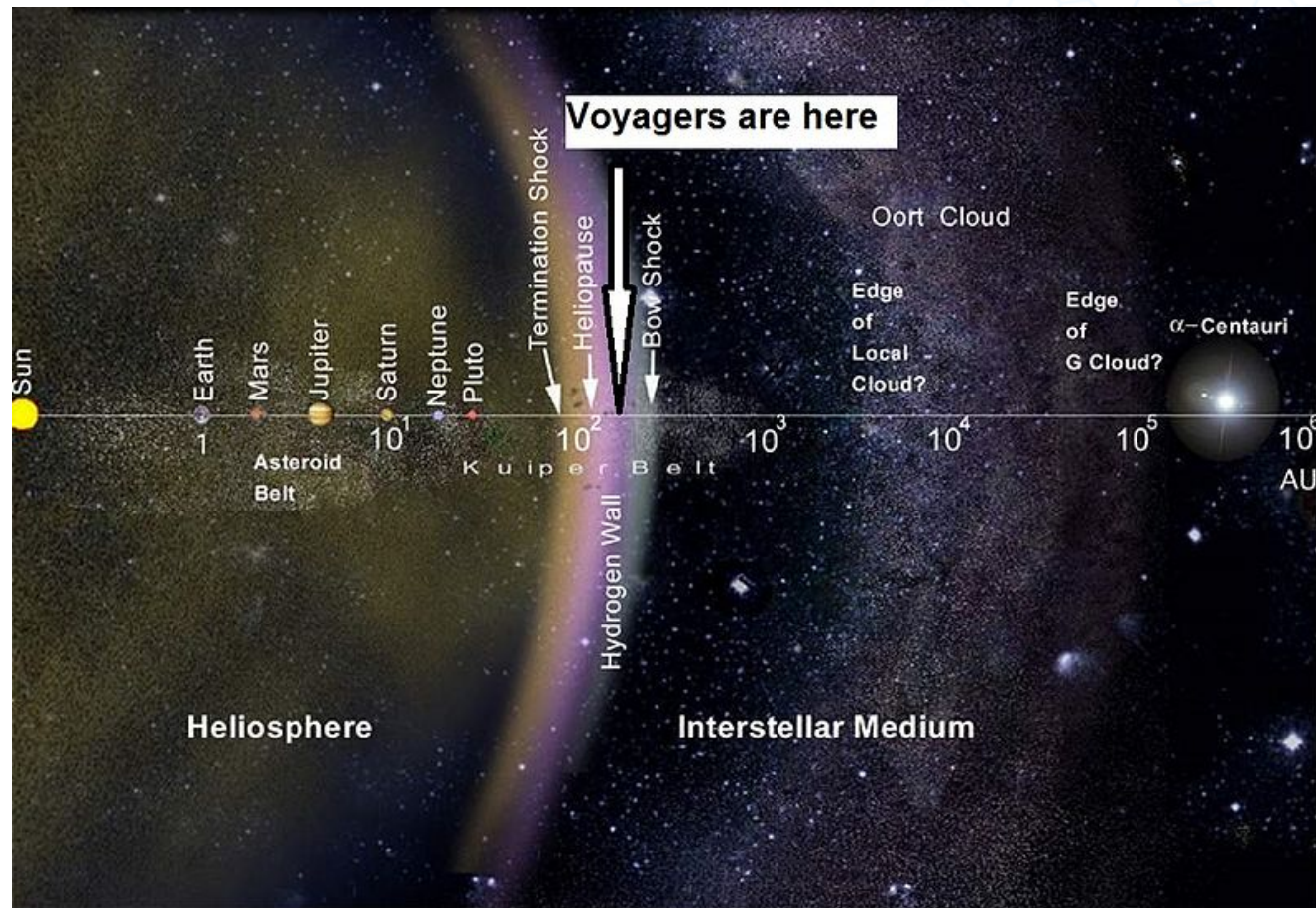
Heliosphere

1993, D. Gurnett reported the first evidence of the heliopause based on the kHz radio emissions coming from the heliopause and detected by the Voyager 1 and 2 spacecraft.



Voyager 1 and 2 crossed the termination shock

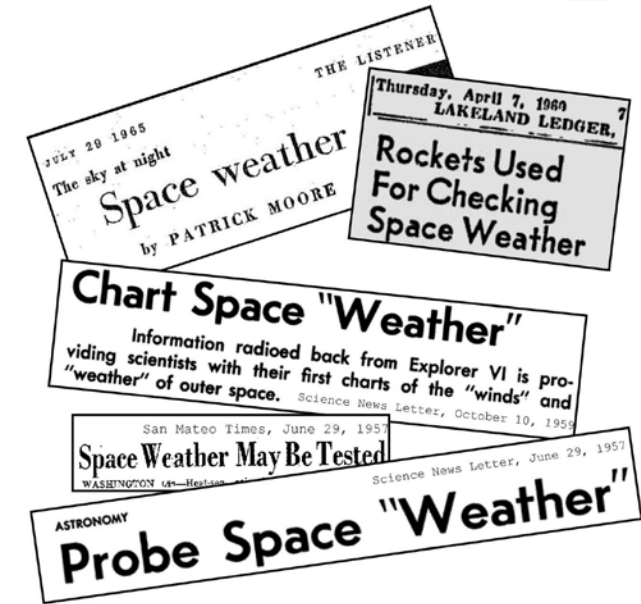
The twin Voyager 1 and 2 spacecrafts are still exploring the border of solar system.



3. The Space Weather Concept and Current Programs

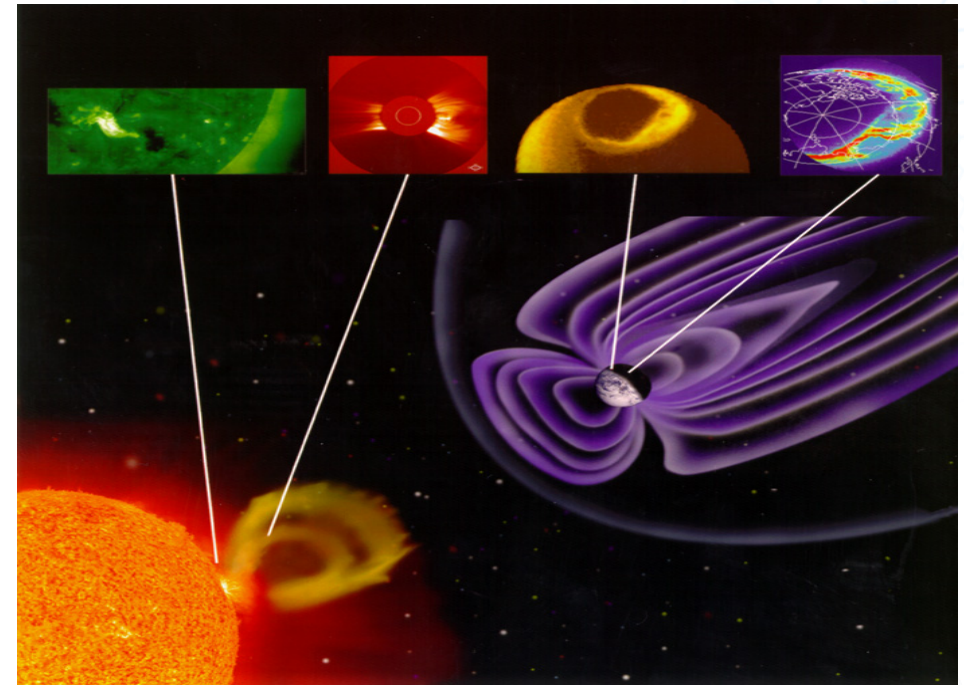
When is the term 'Space Weather' come into being?

The term space weather was first used in the 1950s and came into common usage in the 1990s in the National Space Weather Plan of US.



The definition of Space Weather

Space Weather is the conditions on the sun and in the solar wind, magnetosphere, ionosphere, and thermosphere that can influence the performance and reliability of space-borne and ground-based technological systems, and can endanger human life or health.



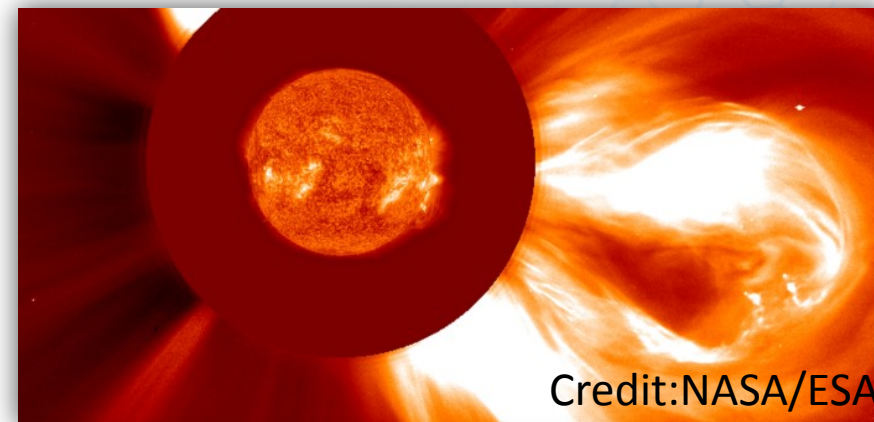
[US National Space Weather Plan]

Weather



Hurricanes and Tornadoes

Space Weather



Credit: NASA/ESA

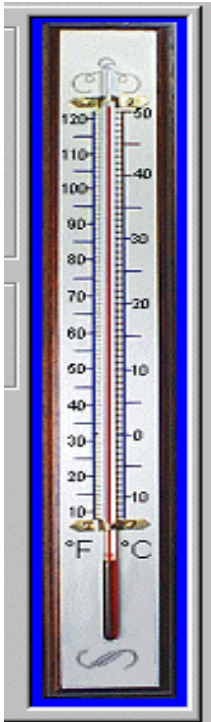


Credit: ISS

Solar Flares and Coronal Mass Ejection, Aurora

Monitor and Measure

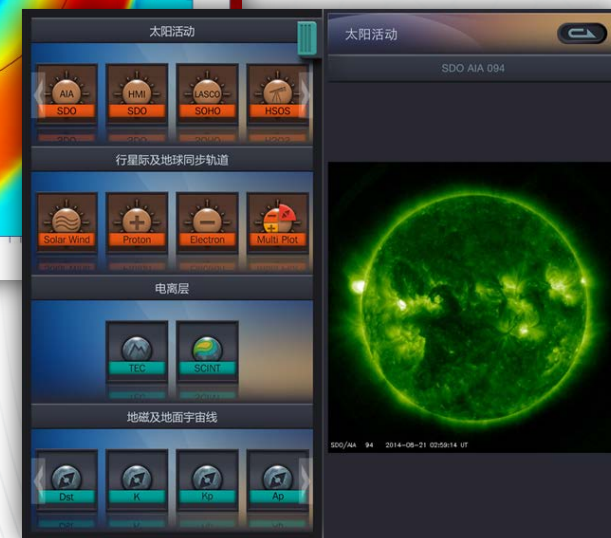
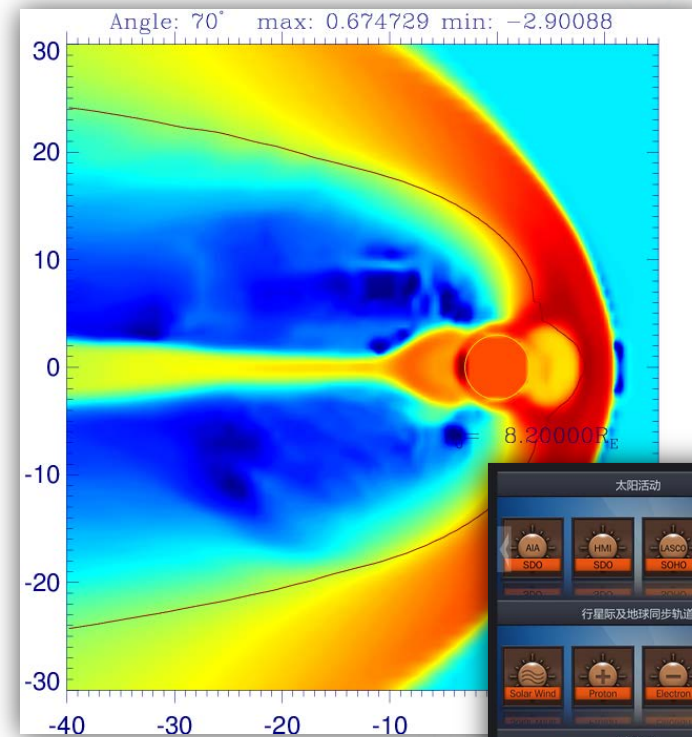
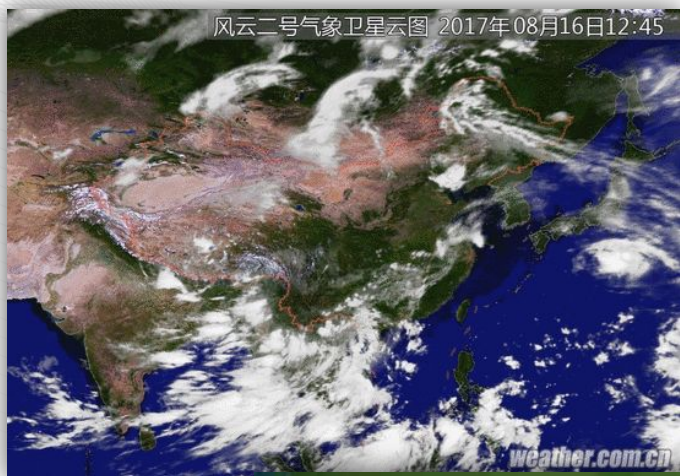
Thermometer



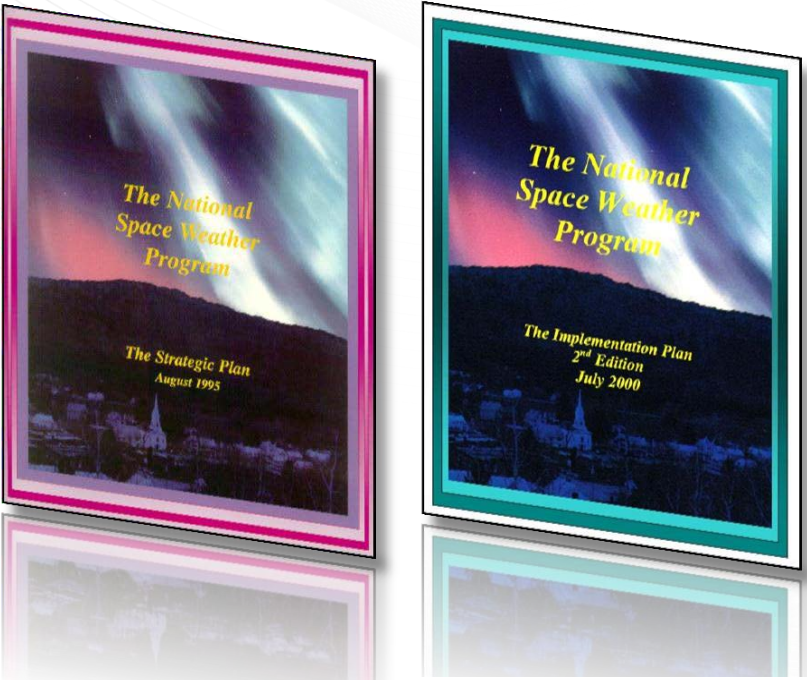
Energetic Particle Sensor



Services



National Space Weather Program (NSWP) — USA



The National Space
Weather Program
The Strategic Plan of US

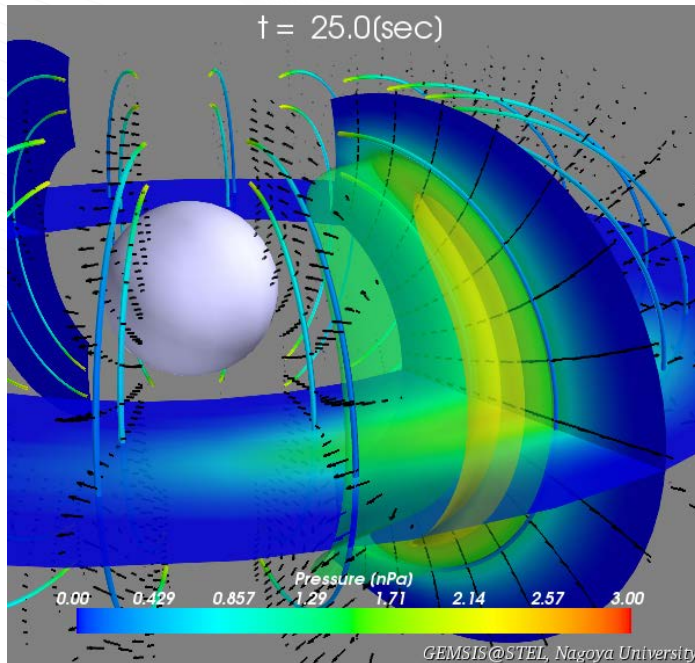
The **National Space Weather Program** (NSWP) established in 1995 with publication of Strategic Plan.

- Pulled federal community together
- Set a vision for the future



Space Weather
Team in U.S.

George L. Siscoe

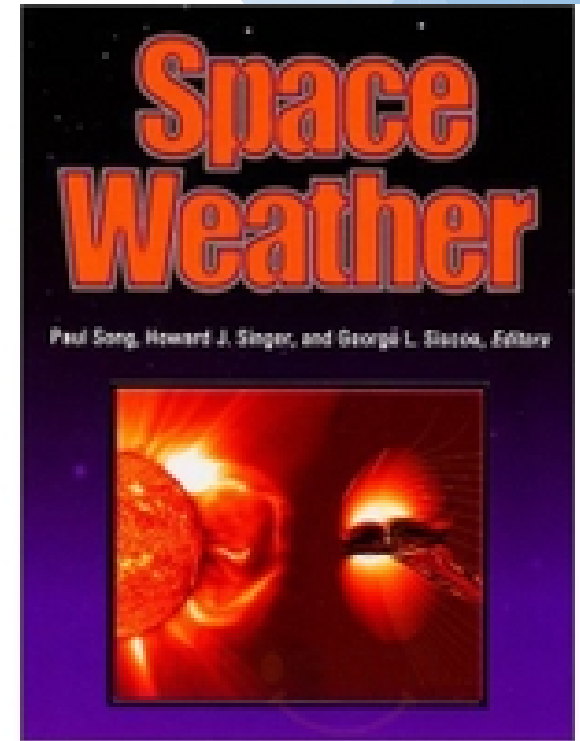


- George Siscoe initiated the Geospace Environment Model plan (GEM).
- It is this plan that led to the National Space Weather Plan established by several Departments of US.



George L. Siscoe

- George Siscoe published a series of impactful papers concerned about space weather, and he was one of the editors of the book "Space Weather".
- He also was the first editor-in-chief of the "Space Weather" journal.



The Development of European Space Weather

1996: ESA Round Table on Space Weather.

1998: First ESA Space Weather Workshop.

1999-2001: ESA feasibility study on a Space Weather Programme.

2000: Setting up of Space Weather Working Team.

2003: ESA Space Weather pilot-project formally starts.

European Space Weather Program focused on monitoring conditions at the Sun and in the solar wind, and in Earth's magnetosphere, ionosphere and thermosphere, that can affect spaceborne and ground-based infrastructure or endanger human life or health.

Space Weather in Russia

Prof. G. A. Zherebtsov made a significant progress Space Weather especially focus on the incoherent scatter radar, ionosphere, global climate change, earth observation, solar observation in Russian.

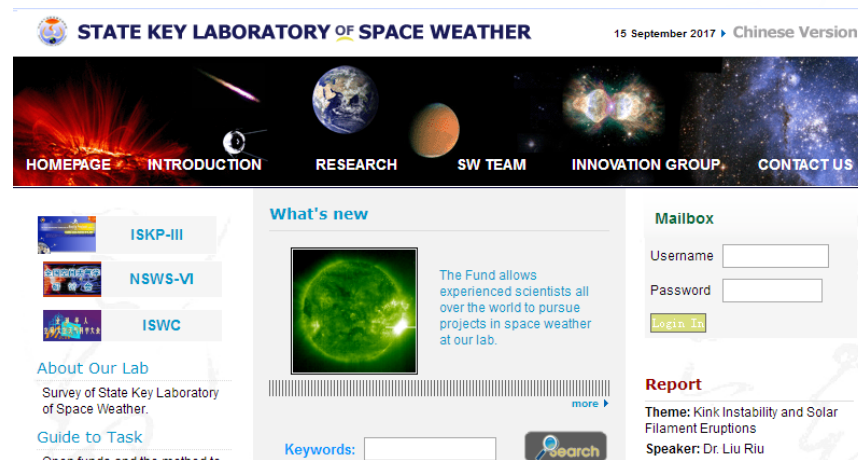
Jointly with ISTP, IKI and IZIMIRAN are also proposed may observation and missions for space weather studies.



Space Weather in China

China established its own Space Weather Laboratory, Space Weather Prediction Service and a roadmap to 2050.

The State Key Laboratory of Space Weather was founded in 2006, which is the first state key laboratory approved by the Ministry of Science and Technology in the field of space physics in China.



Academician Wei Fengsi

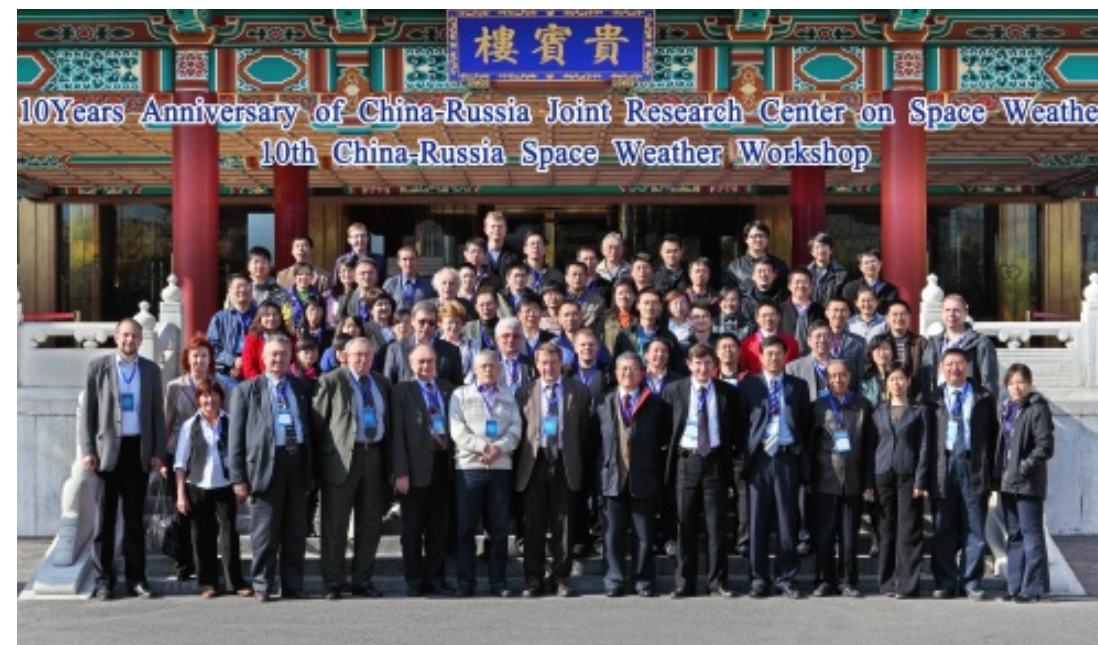
- **Academician Wei Fengsi** initiated 'Space Weather' concept in China. He contributed to the establishment of the State Key Laboratory of Space Weather.
- He dedicated in promoting Chinese Meridian Project.
- He proposed the International Space Weather Conference which enhance the global cooperation in space weather.



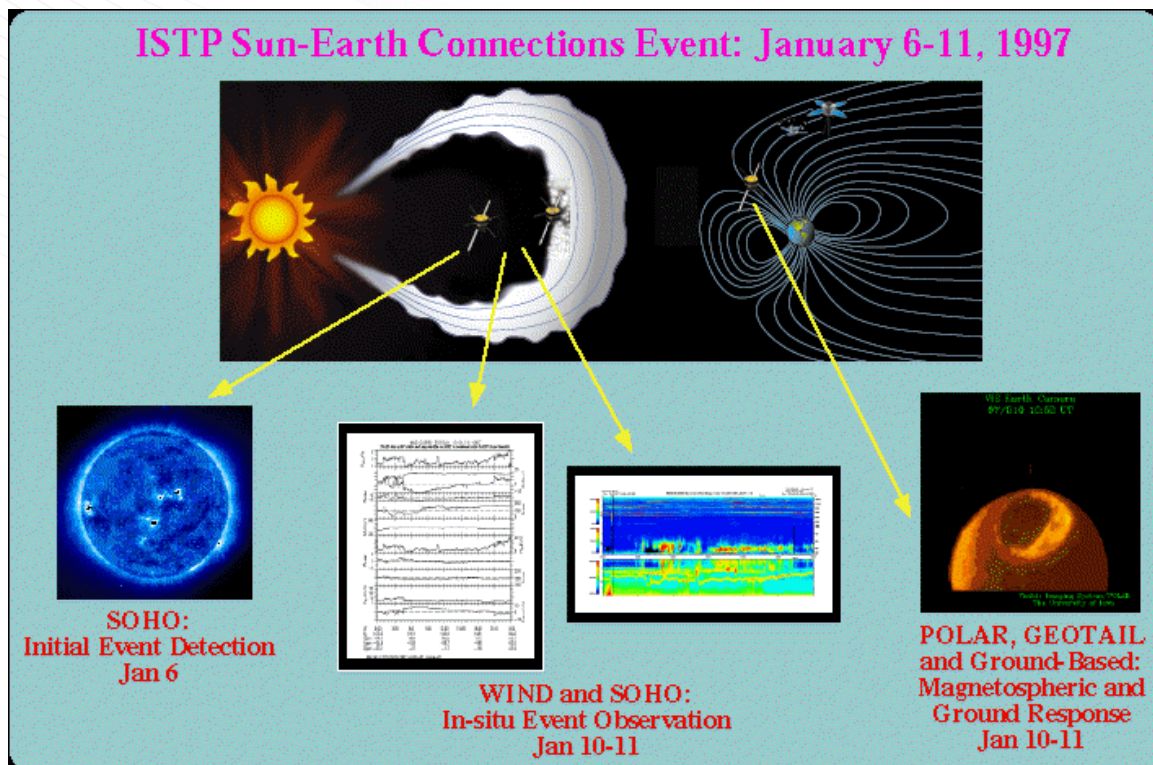
China and Russian's contribution to Space Weather

Sino-Russia Joint Space Weather Research Center

For more than a decade, we have held bilateral seminars 11 times (once a year)

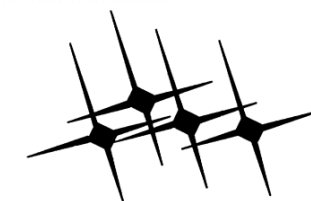


International Solar-Terrestrial Physics



The **International Solar Terrestrial Physics** (ISTP) Program is a large, multi-national program involving three space agencies and up to eight spacecraft.

IACG for Space Sciences was formed in 1981 and until 1986 coordinated the six space missions to Halley's Comet.



Inter-Agency Consultative Group
for Space Science (IACG)

Living With a Star

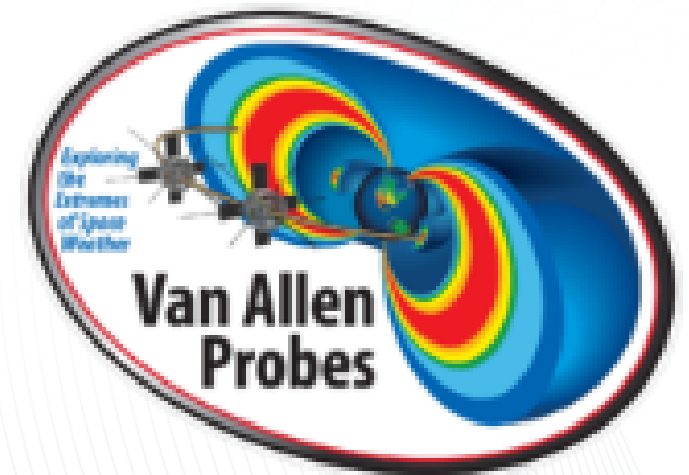
Living With a Star (LWS) is a NASA scientific program to study those aspects of the connected Sun-Earth system that directly affect life and society. LWS is a crosscutting initiative with goals and objectives relevant to NASA's Exploration Initiative, as well as to NASA's Strategic Enterprises. The program is managed by the Heliophysics Division of NASA's Science Mission Directorate.

Operating Missions:

- Van Allen Probes (Radiation Belt Storm Probes-RBSP) Mission
- Solar Dynamics Observatory (SDO) Mission
- Stereo A and B

Missions in Development

- Space Environment Testbeds (SET)
- Solar Orbiter Collaboration
- Solar Probe Plus

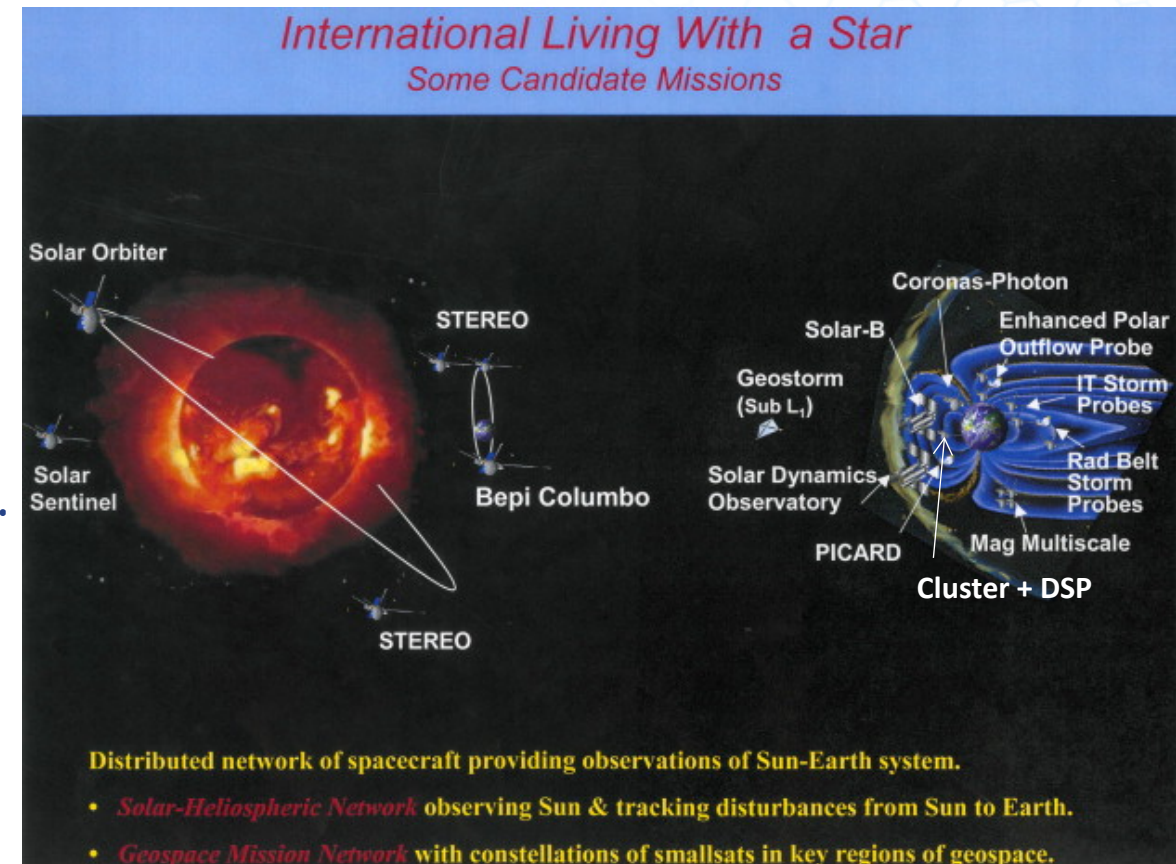


International Living With a Star

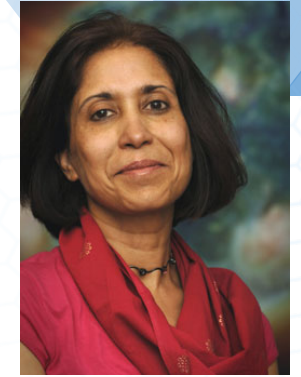


MISSION:

Stimulate, strengthen, and coordinate space research to understand the governing processes of the connected Sun-Earth System as an integrated entity.

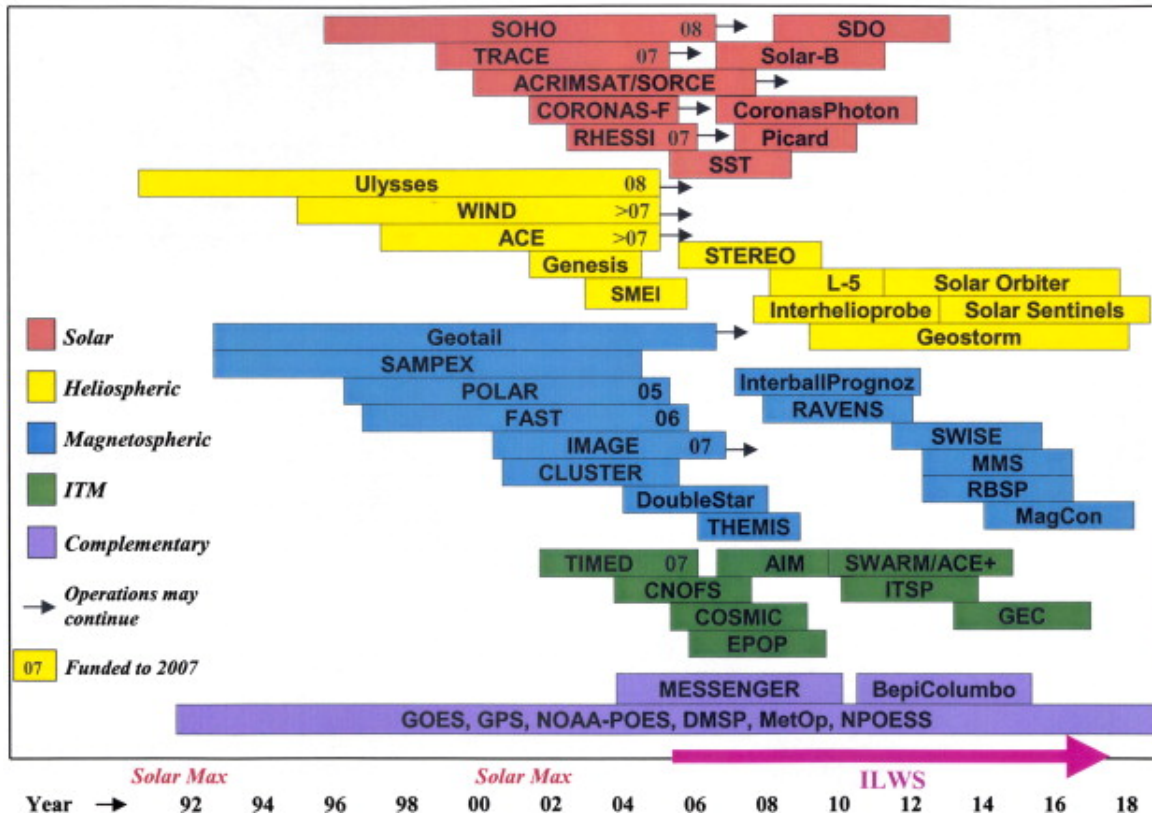


International Living With a Star



Madhulika Guhathakurta

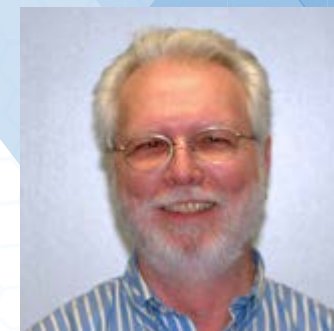
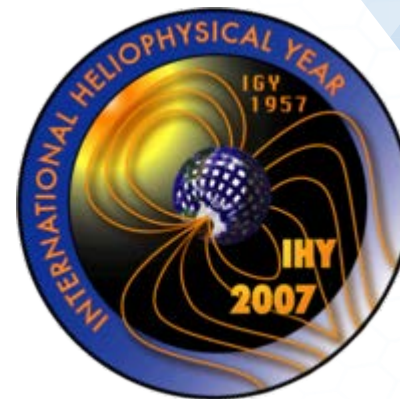
Present Solar-Terrestrial Missions & “First Order” ILWS Missions



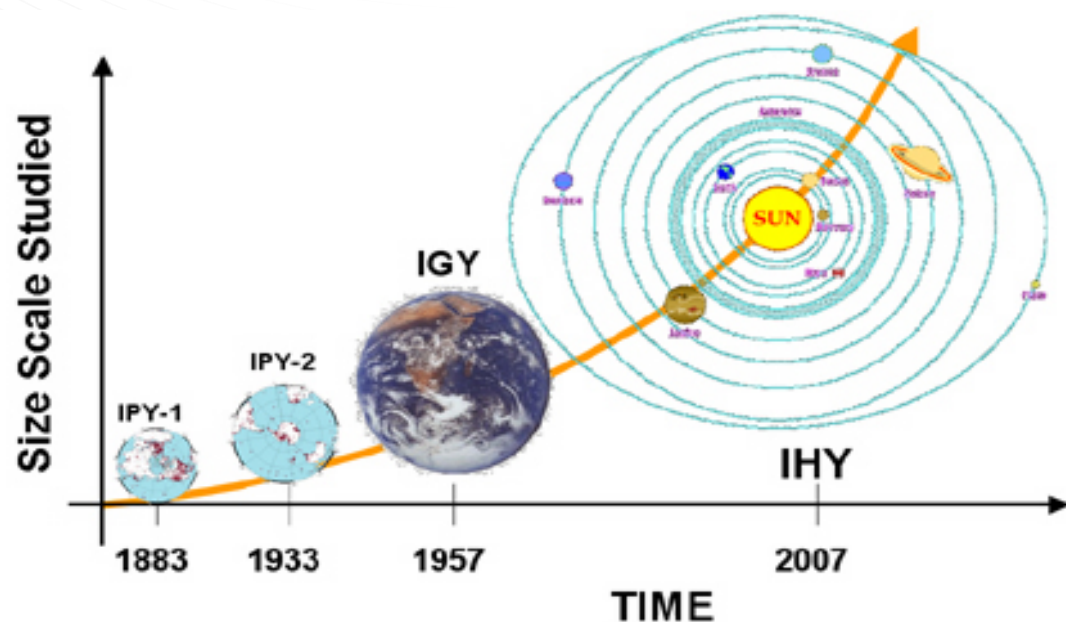
OBJECTIVES:

- Study of the Sun-Earth connected system and the effects which influence life and society.
- Collaboration among potential partners in solar-terrestrial space missions.
- Synergistic coordination of international research in solar-terrestrial studies, including all relevant data sources as well as theory and modeling.
- Effective and user driven access to all data, results, and value-added products.

International Heliophysical Year



Joseph Davila

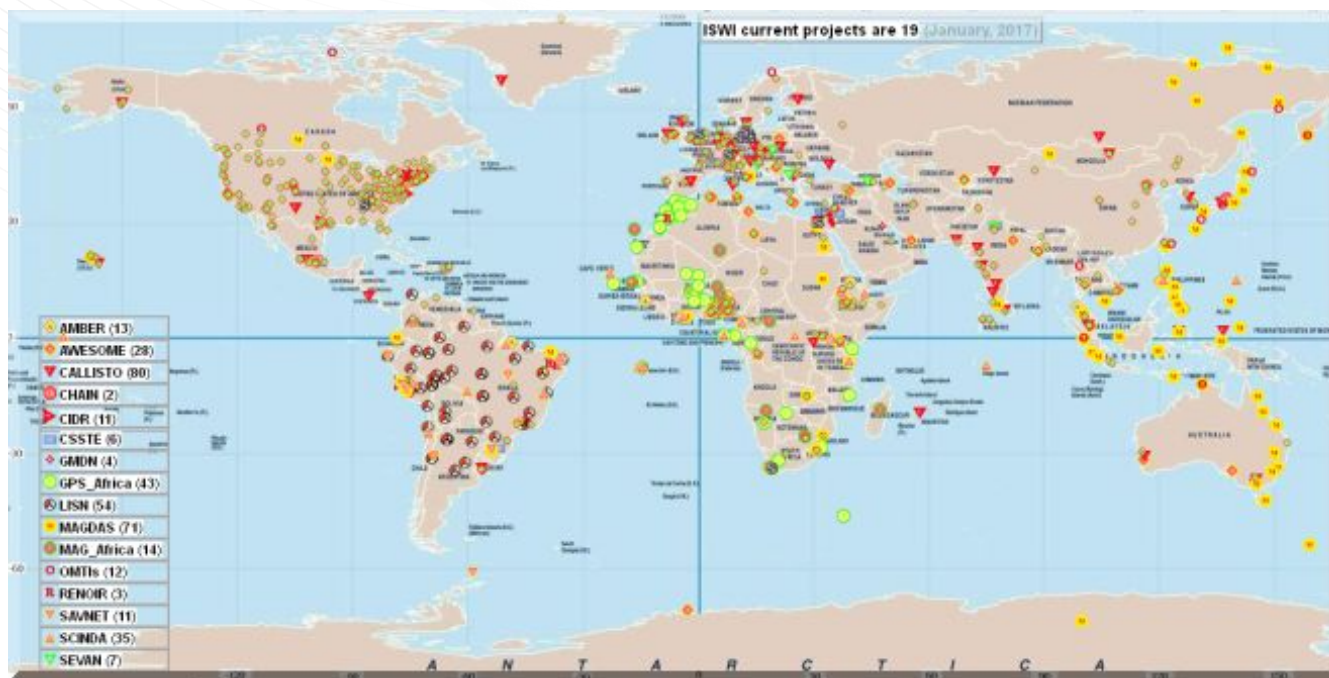


The **International Heliophysical Year** is a UN-sponsored scientifically driven international program of scientific collaboration to understand external drivers of planetary environments and universal processes in solar-terrestrial-planetary-heliospheric physics.

The International Space Weather Initiative



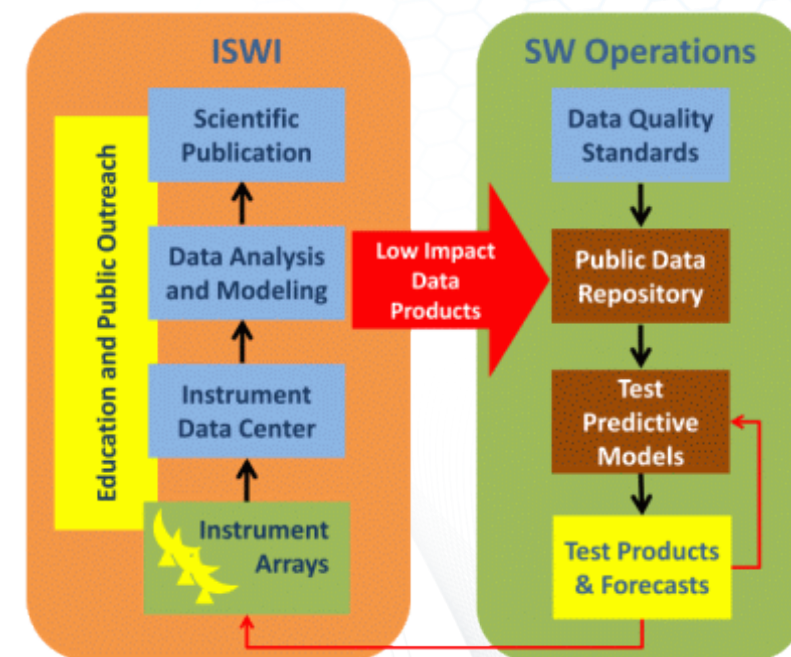
ISWI is a program of international cooperation to advance the space weather science by a combination of instrument deployment, analysis and interpretation of space weather data from the deployed instruments in conjunction with space data, and communicate the results to the public and students.





The Goals of ISWI

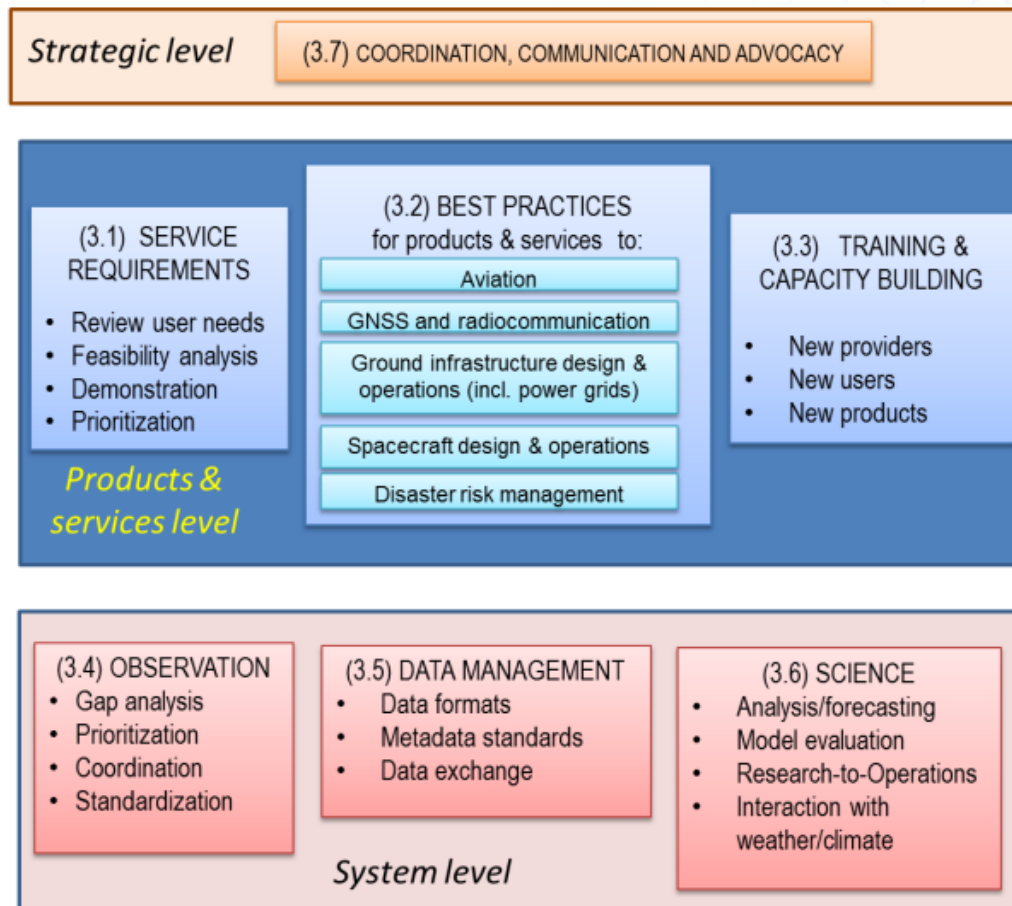
- To develop the scientific insight necessary to understand the science,
- To reconstruct and forecast near-Earth space weather.
- These include instrumentation, data analysis, modeling, education, training, and public outreach.



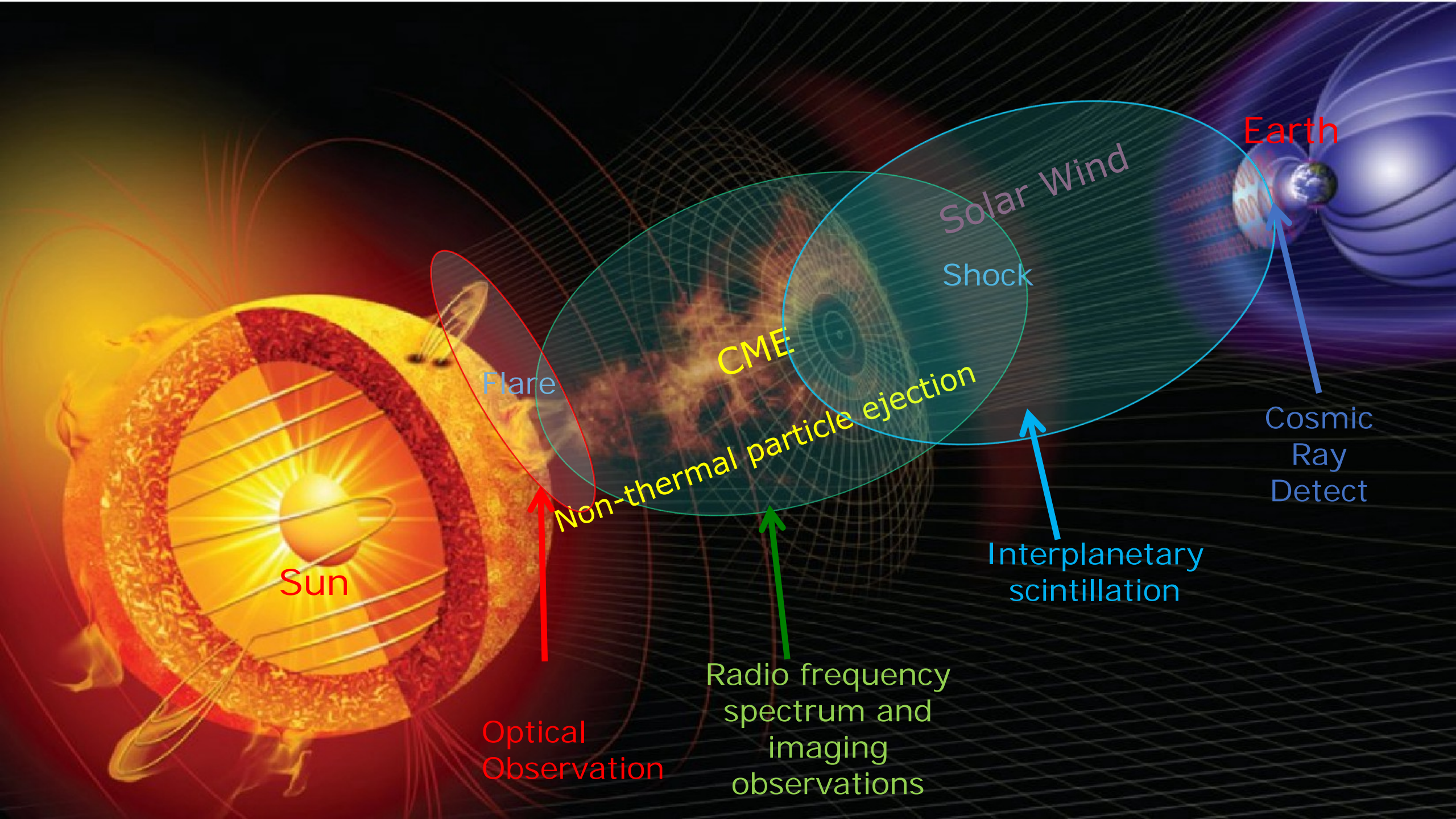


Space Weather program of WMO

- In May 2010, WMO established the Interprogramme Coordination Team on Space Weather (ICTSW) with a mandate to support Space Weather observation, data exchange, product and services delivery, and operational applications.
- As of May 2016, ICTSW involves experts from 26 different countries and 7 international organizations.



4. Looking at the Future Space Weather Programs



Sun

Flare

Optical
Observation

CME

Non-thermal particle ejection

Radio frequency
spectrum and
imaging
observations

Solar Wind

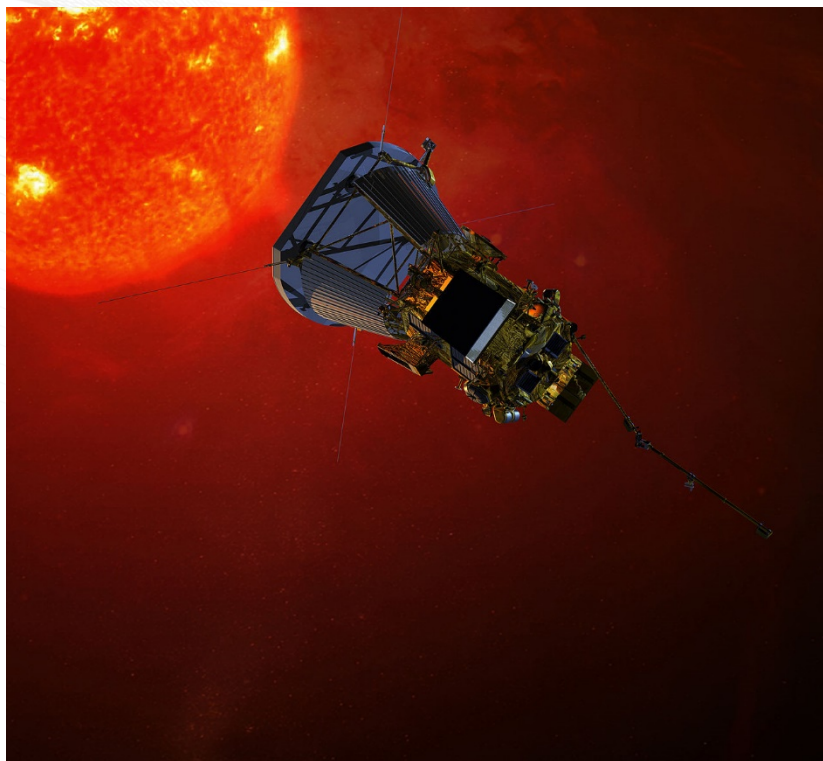
Shock

Interplanetary
scintillation

Earth

Cosmic
Ray
Detect

Parker Solar Probe (Solar Probe +)



- Parker Solar Probe (previously Solar Probe, Solar Probe Plus, or Solar Probe+) is a planned NASA robotic spacecraft to **probe the outer corona of the Sun**. On May 31, 2017 the probe was renamed after solar astrophysicist Eugene Parker. This was the first time a NASA spacecraft was named after a living person
- It will approach to within **8.5 solar radii** (5.9 million kilometers or 3.67 million miles) to the '**surface**' (photosphere) of the Sun

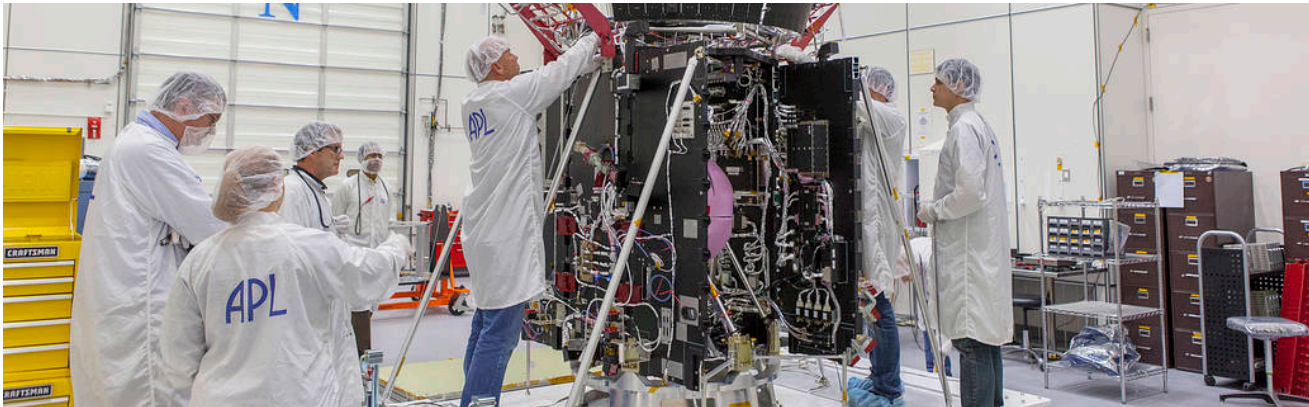
Perihelion	6.0 million km; 0.040 AU (3.7 million mi)
Aphelion	109.3 million km; 0.730 AU (67.9 million mi)
Inclination	3.4°
Period	88 days



Parker Solar Probe (Solar Probe +)

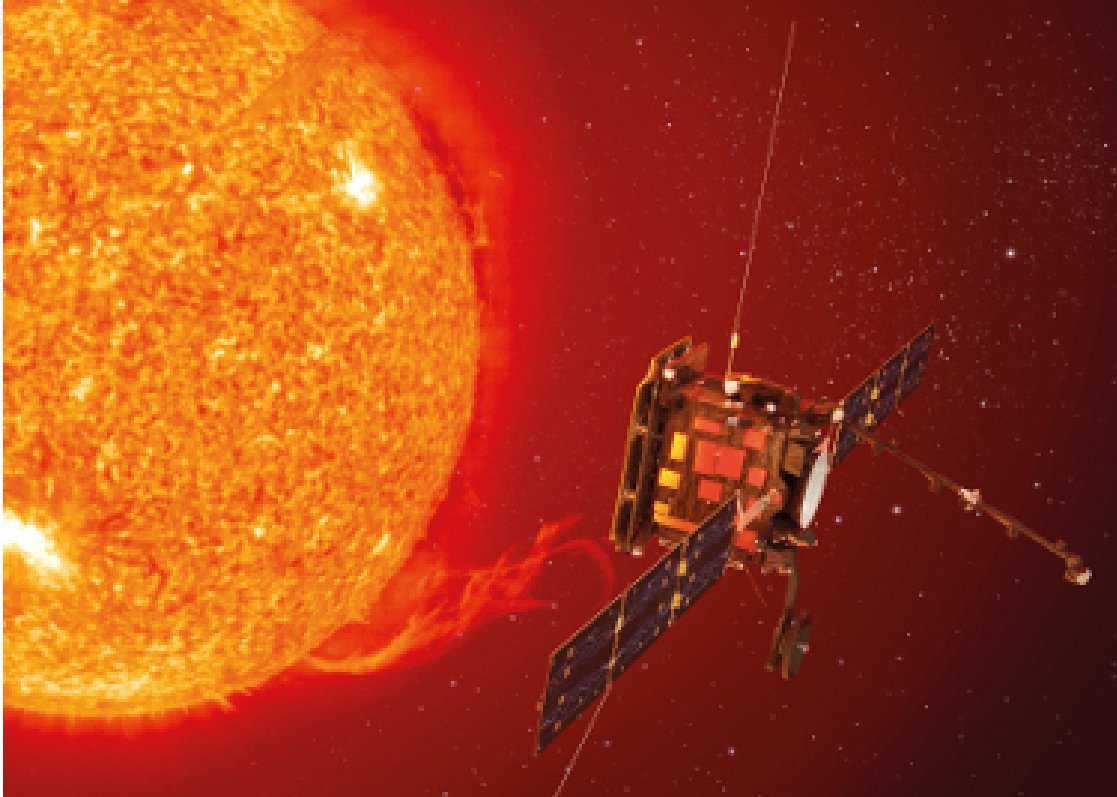
Scientific goals

- Determine the structure and dynamics of the magnetic fields at the sources of solar wind.
- Trace the flow of energy that heats the corona and accelerates the solar wind.
- Determine what mechanisms accelerate and transport energetic particles.
- Explore dusty plasma near the Sun and its influence on solar wind and energetic particle formation.



- The project was announced as a new mission start in the fiscal 2009 budget year.
- The launch date has since been pushed back to 2018.

Solar Orbiter observations of the polar regions of the Sun, which is difficult



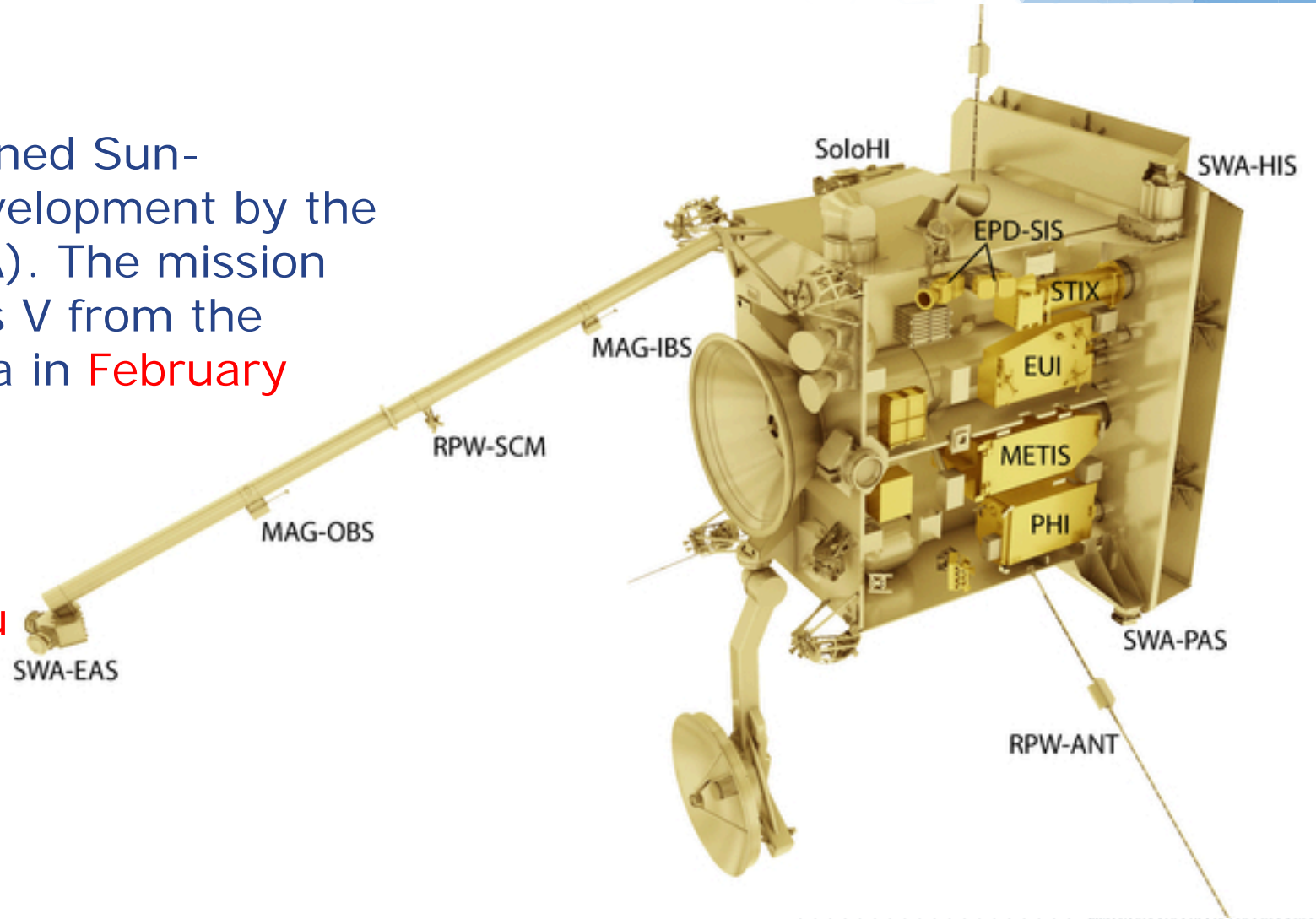
Perihelion	0.28 AU
Aphelion	0.8-0.9 AU
Inclination	0-34 degrees
Period	150 days

- SolO is intended to perform detailed measurements of the inner heliosphere and nascent solar wind, and perform close to do from Earth, both serving to answer the question 'How does the Sun create and control the heliosphere?'
- The Solar Orbiter will make observations of the Sun from an eccentric orbit moving as close as ~ 60 solar radii (RS), or 0.284 astronomical units (AU), placing it inside Mercury's perihelion of 0.3075 AU and providing it with the closest ever views of the Sun.

Solar Orbiter

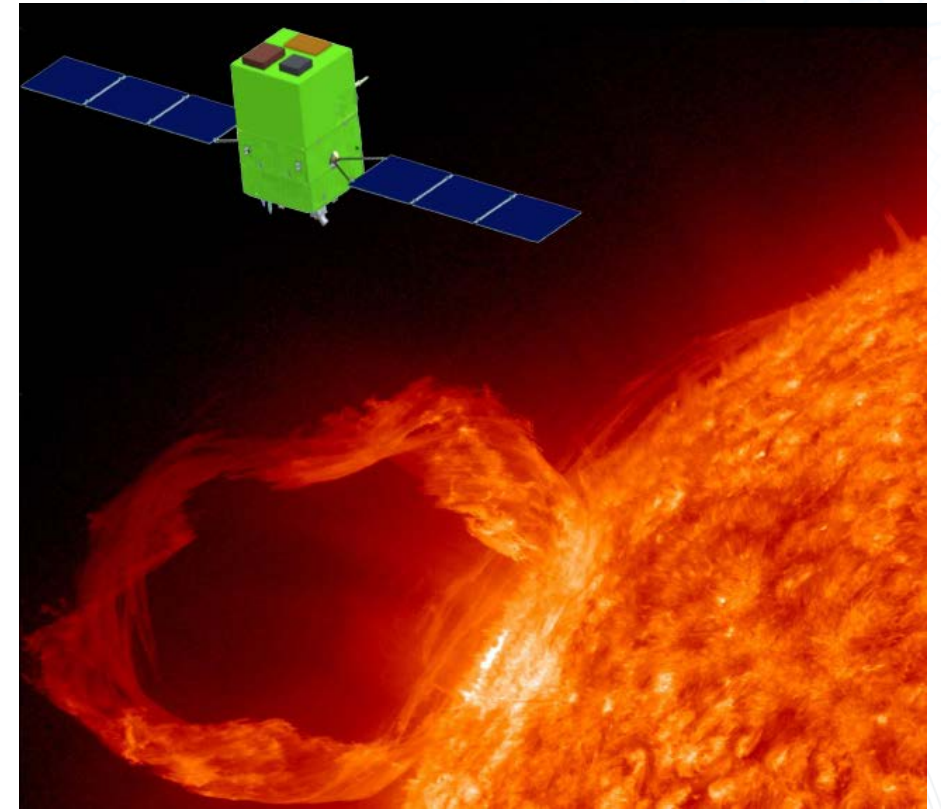
Solar Orbiter (SoIO) is a planned Sun-observing satellite, under development by the European Space Agency (ESA). The mission will be launched with an Atlas V from the Cape Canaveral AFS in Florida in **February 2019**.

The Solar Orbiter payload accommodates a set of **in situ** and a set of **remote-sensing** instruments, with a total payload mass of 180 kg.



Major Scientific Objectives for ASO-S (Advanced Solar Observatory in Space)

- **Simultaneously observe** the full disc vector magnetic field, non-thermal images of hard X-rays, and initiation of CME
- **Understand** the causality between magnetic field and flares, magnetic field and CMEs, flares and CMEs



ASO-S Orbits

■ Full-Disc Vector Magnetograph (FMG)

to observe full-disc vector magnetic field of photosphere with a high time resolution. Heritage: SST/DSO as well as HSMT

■ Lyman-alpha Solar Telescope(LST)

to observe Lyman-alpha disc image + inner corona
Heritage: SMESE and some pre-studies

■ Hard X-ray Imager (HXI)

to spectrally image the Sun in 30-300 keV
Heritage: SZ-02/Chang'E/SMESE and some pre-studies

Solar Synchronous

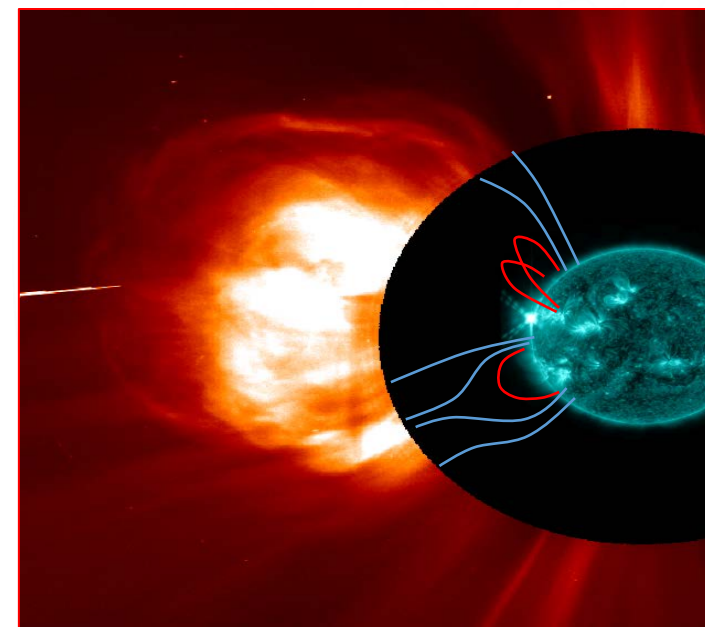
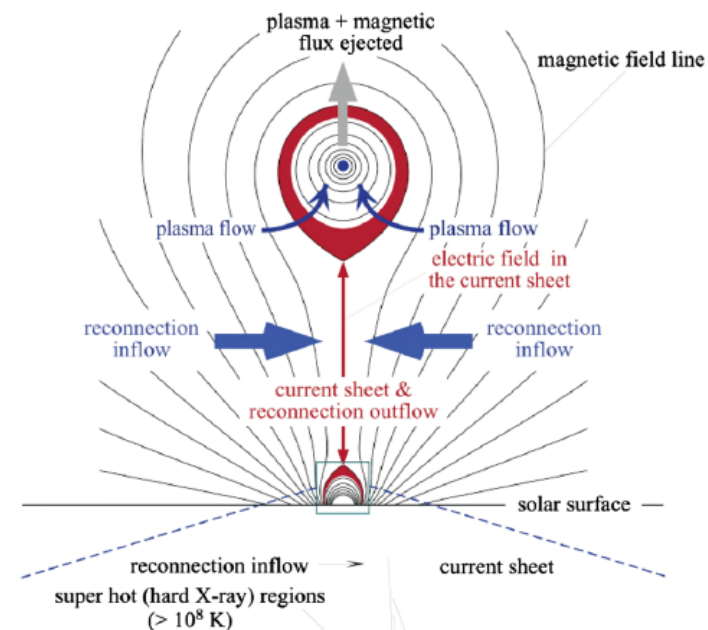
Attitude: 700 - 750 km

Inclination: 98.27 °

Launch date: 2022

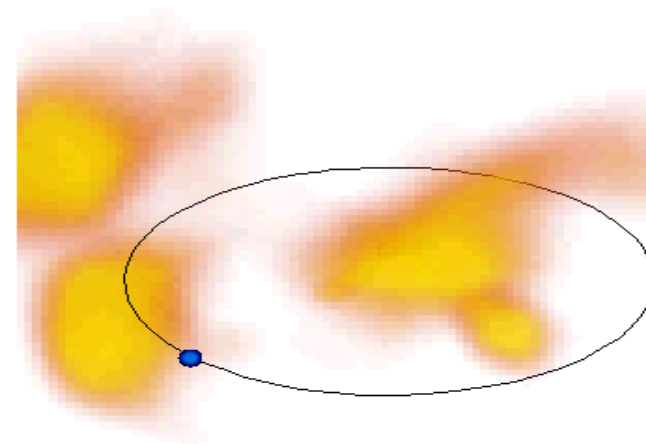


中国科学院国家空间科学中心
National Space Science Center, CAS

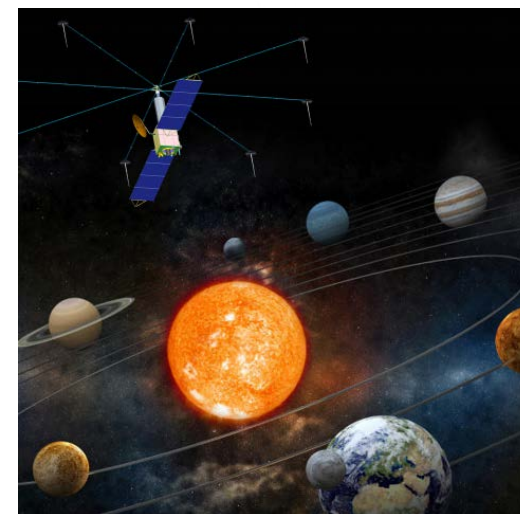


SPORT(Solar Polar ORbit Telescope)

- **Reveal** the complete physical process about the triggering, formation, onset, propagation, and evolution stages of coronal mass ejections (CMEs) in the inner heliosphere, and the ensuing responses of structures and dynamics of the inner heliosphere
- **Discover** the effects of solar high-latitude magnetism on solar storm eruptions in the short term and solar cycle variations in the long term.
- **Investigate** the origin and properties of the fast solar wind
- **Understand** the acceleration, transport and distribution of energetic particles in the corona and heliosphere, and their causal relations with solar eruptions



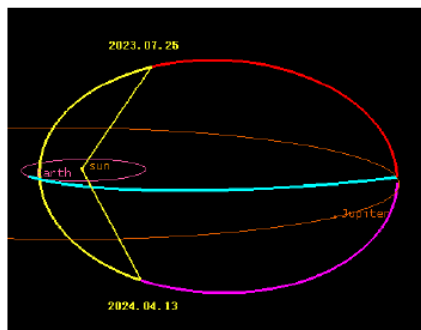
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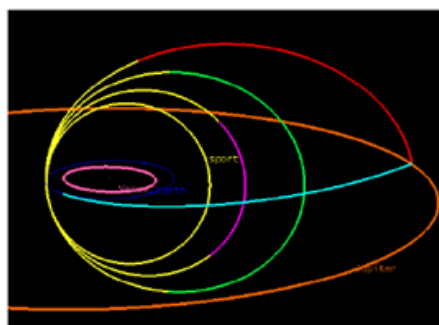
SPORT

Solar polar orbit (inclination > 60°),
periapsis to the sun: 0.5-1AU

- ✓ Time for Imaging Observation
 - Flying over the solar polar regions, with the SPORT-sun distance within 2AU
 - During the next solar maximum
- ✓ Use multi-gravity assists (further Venus/Earth swing-by) to increase the

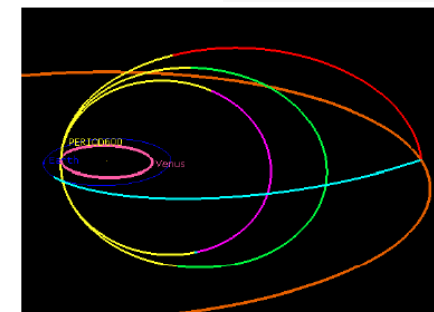


Imaging observation: 17 months
2023.07.25-2024.04.13
2028.06.06-2029.02.25

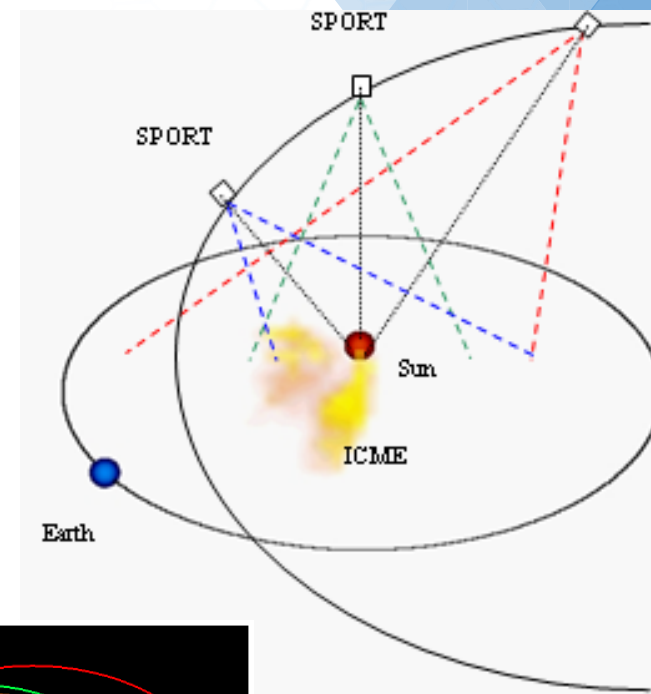


Imaging observation: 45 months
2023.09.25-2024.07.30
2026.08.25-2027.10.04
2028.06.15-2030.03.20

NSSC 中国科学院国家空间科学中心
National Space Science Center, CAS



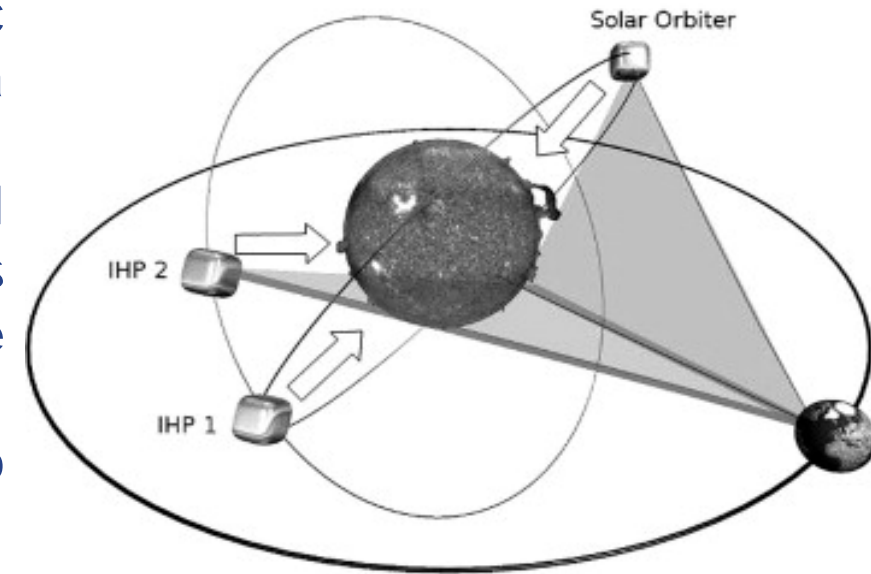
Time of observation: 32 months
2023.07.27-2024.05.02
2026.08.06-2027.06.24
2028.09.05-2029.08.20



Interhelioprobe

- The missions have been designed by the Russian and European Space Agencies and are aimed at studying the polar and equatorial regions from high heliolatitudes.
- Their primary task is to investigate the polar magnetic fields, plasma motions and solar dynamo, ecliptic corona and heliolatitudinal structure of mass ejections, mechanisms of corona heating and solar wind acceleration, triggering of solar flares and mass ejections, mechanisms of particle acceleration in the Sun and the heliosphere, solar wind sources in the Sun, and the relation of solar transient phenomena to variations in the heliosphere

Interhelioprobe/Solar Orbiter Cooperation



Interhelioprobe

Payload

SOLAR INSTRUMENTATION-10

Optical photometer

Magnetograph

Chemical Composition Analyzer

EUV Imager-Spectrometer

Coronagraph

X-ray Imager

Heliospheric Imager

X-ray Polarimeter

Gamma-Spectrometers - 2

HELIOSPHERIC INSTRUMENTATION-8

Solar Wind Ion Analyzer

Solar Wind Electron Analyzer

Solar Wind Plasma Analyzer

Energetic Particle Telescope

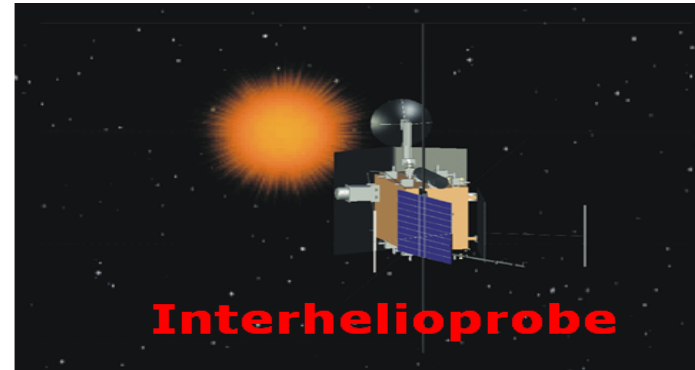
Neutron Detector

Magnetic Wave Complex

Magnetometer

Radio Spectrometer Detector

• **Mass: 120-180 kg** • **Power: 120 W** • **Telemetry: 1 Gb/day**



Heritage:

CORONAS-I (1994-2001)

INTERBALL (1995-2001)

CORONAS-F (2001-2005)

CORONAS-PHOTON (2009)

RESONANS (2014)

Major Scientific Objectives for SMILE

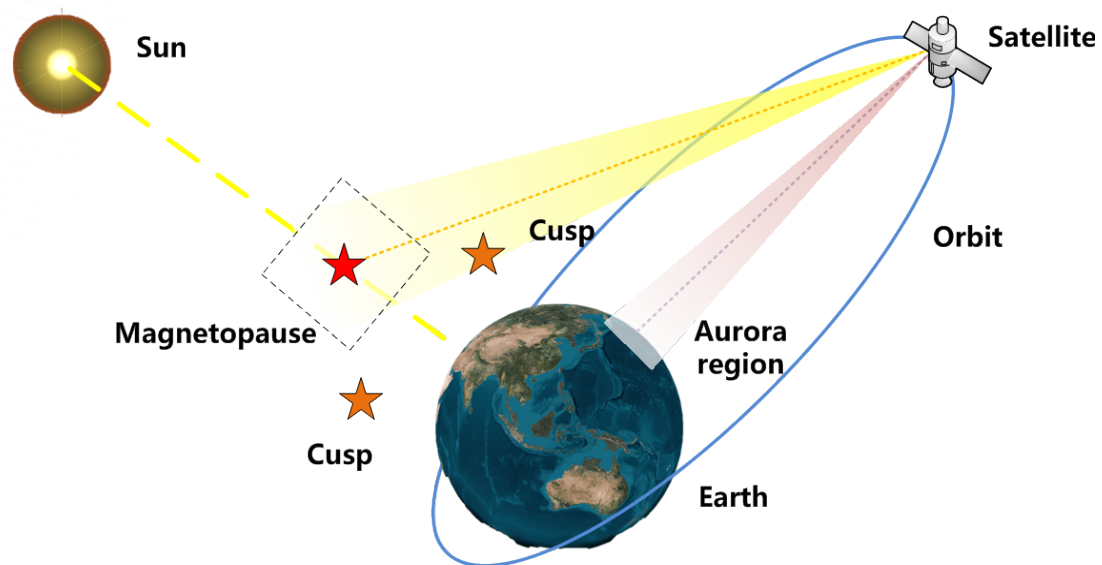
- Investigate the dynamic response of the Earth's magnetosphere to the solar wind impact in a unique and global manner
- Combine X-ray imaging of the dayside magnetosheath and the cusps with simultaneous UV imaging of the northern aurora, while monitoring the solar wind conditions in situ
- Full chain of events that drive Sun-Earth relationships: dayside reconnection / magnetospheric substorm cycle / CME-driven storms



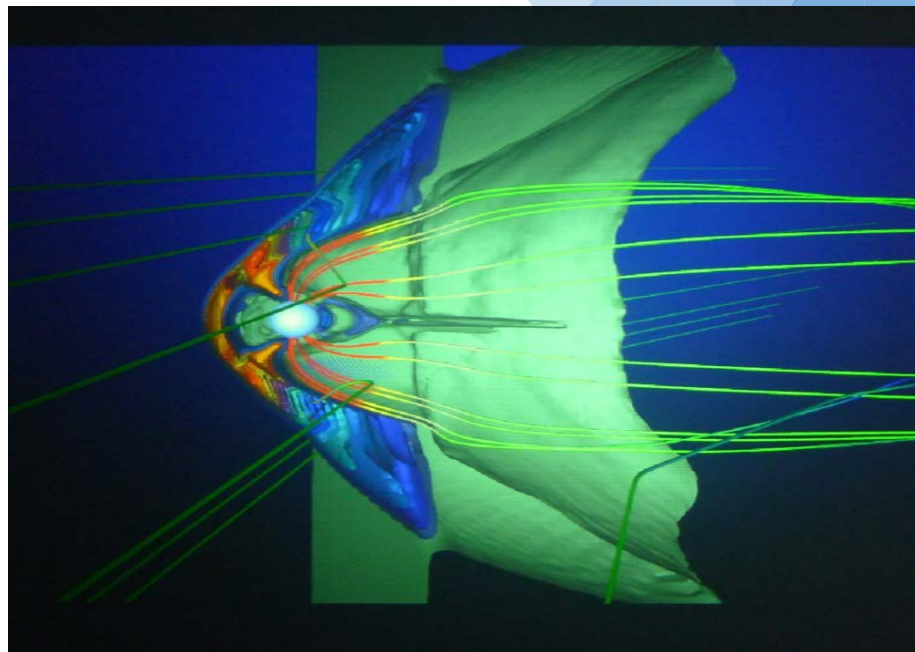
SMILE Orbits

5000km @ perigee

19 RE @ apogee



Launch date: 2022



NSSE

UCL

University of
Leicester

Imperial College
London



UNIVERSITY OF
CALGARY

FINNISH METEOROLOGICAL
INSTITUTE

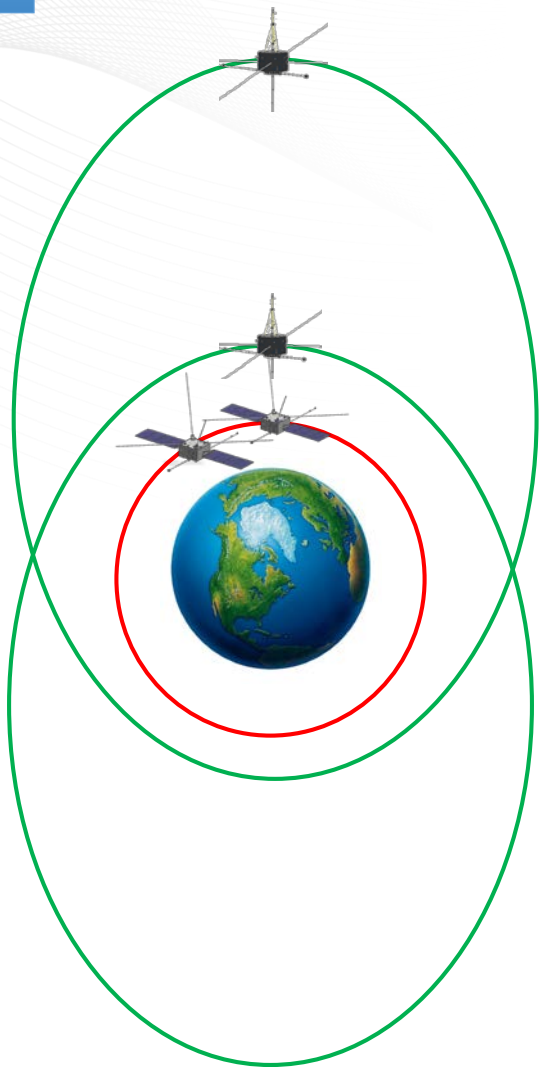
NASA
Goddard



Magnetosphere, Ionosphere and Thermosphere Coupling (MIT)

- **Investigate** the origin of the outflow ions and their acceleration mechanism
- **Understand** the impact of the outflow ions on magnetic storm development
- **Characterize** the ionosphere and thermosphere storm caused by magnetic storm
- **Explore** key mechanisms for the magnetosphere, ionosphere and thermosphere

MIT Orbits



Satellite	Ionosphere/ Thermosphere Satellite-A (ITA)	Ionosphere/ Thermosphere Satellite-B (ITB)	Magnetosphere Satellite-A (MA)	Magnetosphere Satellite-B (MB)
Angle	90°	90°	90°	90°
Perigee Altitude	500 km	500 km	1 Re	1 Re
Apogee Altitude	1500 km	1500 km	7 Re	7 Re

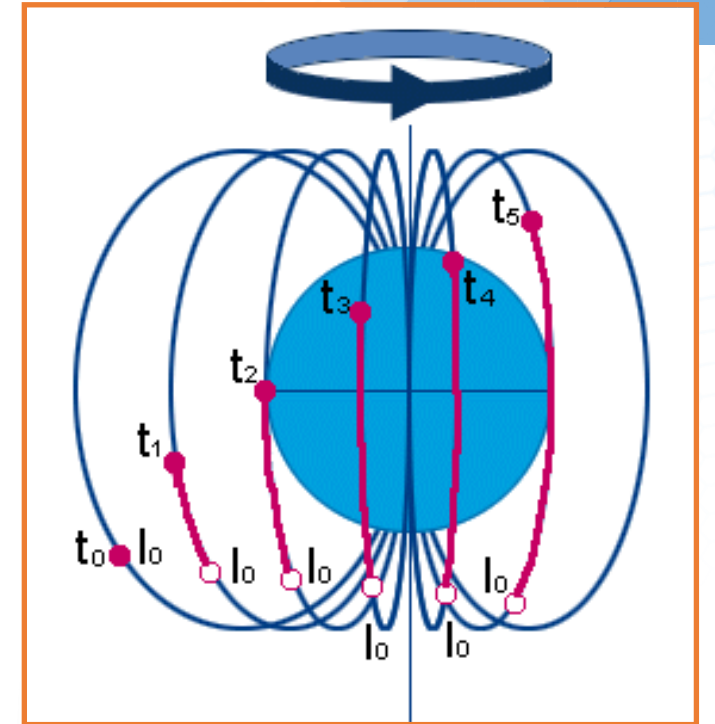
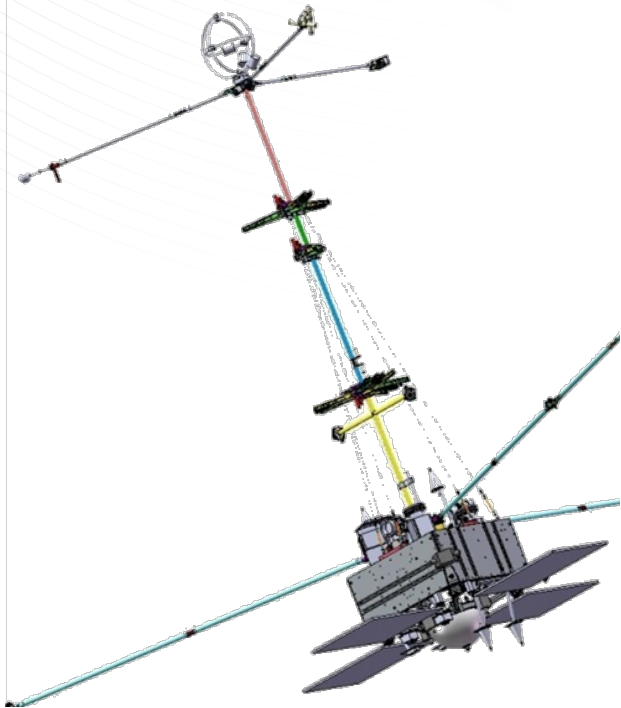
The ratio of MA/MB to ITA/ITB period is 9:1.

Launch date: **2021**

Resonance

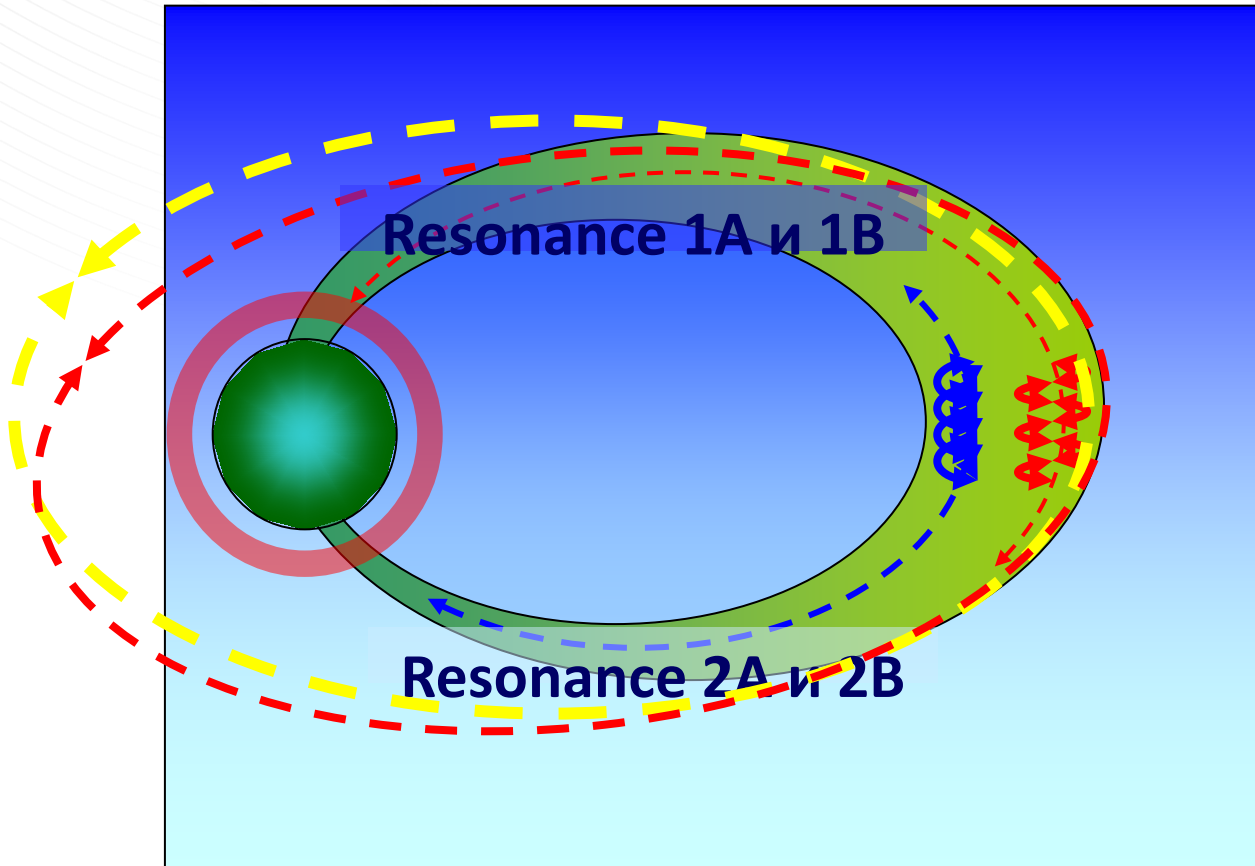
Inner magnetospheric mission

- Space weather
Ring current, outer radiation belt, plasmasphere
- Resonant wave-particle interactions
Magnetospheric cyclotron maser
- Auroral region acceleration
Small-scale active zones, precipitation
- Two pairs of spacecraft
- Magneto-synchronous orbit



Resonance

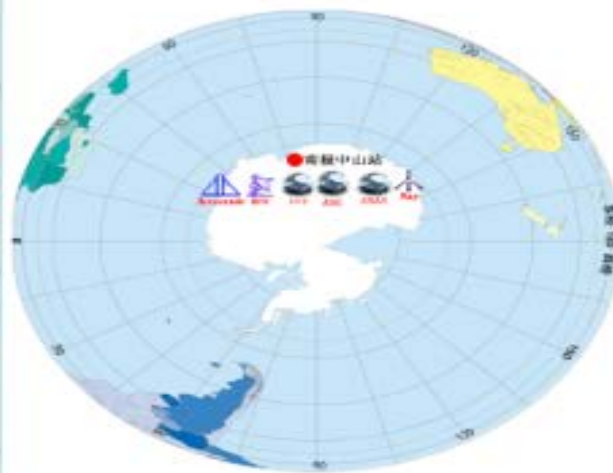
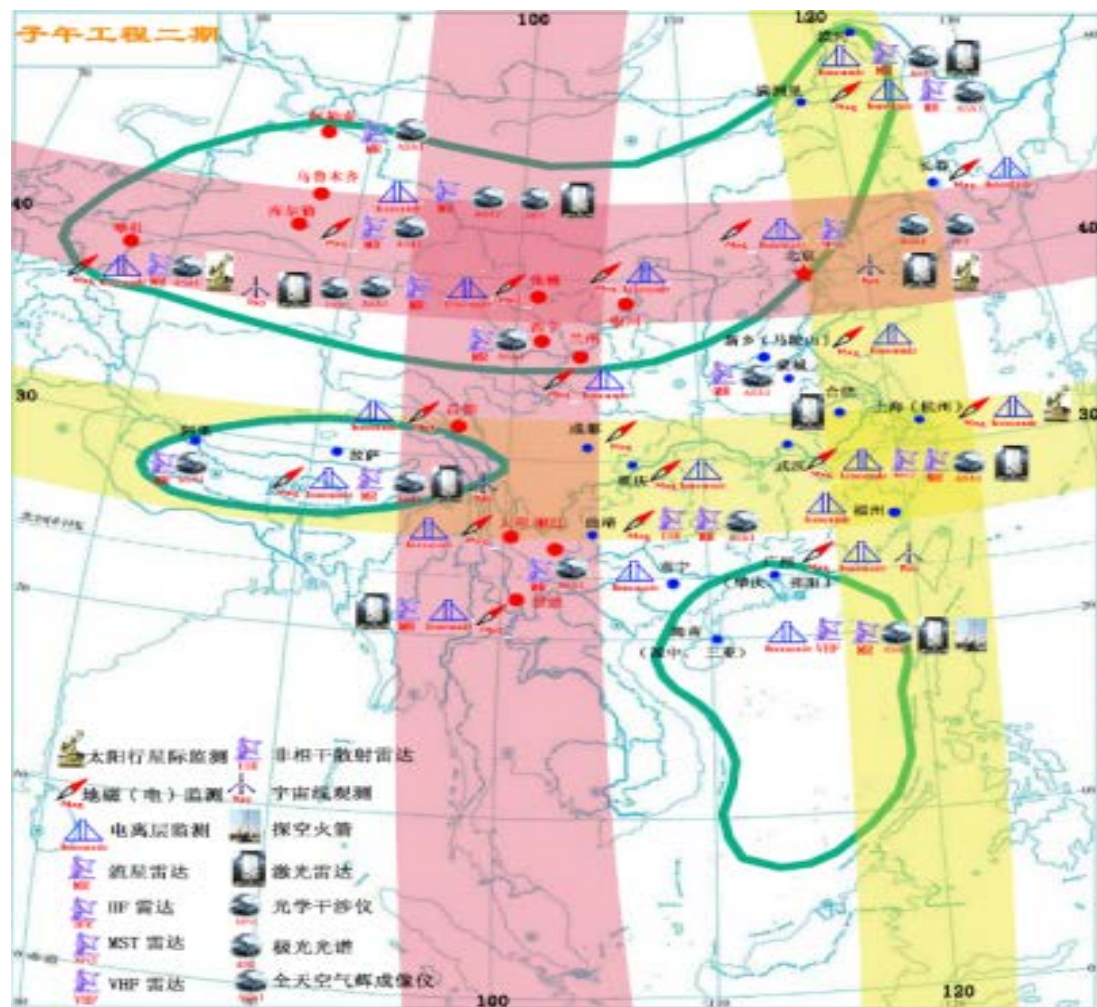
Magnetosynchronous orbit



- Electric and magnetic sensors
Wave analyzer and interferometer
DC – 10 MHz

- Plasma sensors
Cold plasma
Suprathermal plasma
Energetic particles
Relativistic electrons

Chinese Meridian Project II



International Space Weather Meridian Circle Program



International Space Weather Meridian Circle Program (IMCP) connect 120°E and 60°W meridian lines forming chains of ground based observatories to enhance the ability of monitoring space environment worldwide.

China+ Russia+ Australia+ Canada+ Brazil☐
International Meridian Circle Program

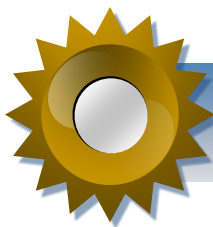


Time Schedule for IMCP

Establishing the
organizational
structure and
objectives

Perfecting the
organizational structure
and carrying out
project cooperation

Forming an
international academic
organization for
multilateral
cooperation



Phase One(2018-2020)

Phase Two(2021-2025)

Phase Three(2025-2030)

Mid-latitude Observation Chain

- Mid-latitude Observation Chain is proposed which start from Japan to Spain which is the **longest (>10,000km) mid-latitude observation chain** on earth.
- It is our opportunity to observe the space environment from the sun to the atmosphere cross **different time zone** at mid-latitude.



Summary

- We have entered the space for 60 years and had a good understanding of the space environment. Space is **not empty, not quiet**, and sometime will **behave badly to us**, the same as the weather on the surface of the earth, therefore we call it – **SPACE WEATHER**
- What we have experienced in space in the past 60 years are less than **6 solar cycles**, each of them are different. The Sun may have much big storms than what we have experienced. So we'd better be prepared.
- To understand the space weather need much more efforts where **new missions** are encouraged and **international collaborations** are very much demanded.

Thank You!

**Congratulations on Human's First Man-made
Satellite Sputnik's 60th Anniversary**

